

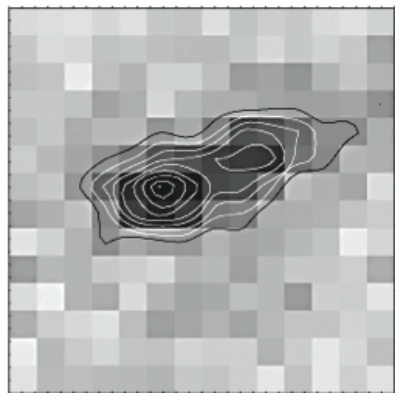
Charting New Physics Territory with Time-Domain Optical Astronomy

Ariel Goobar
The Oskar Klein Centre
Stockholm University

What's new?

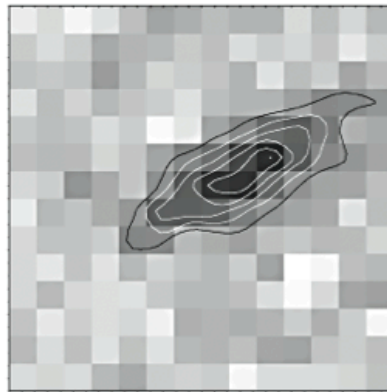
- technological advances!

- Very large field of view: nightly monitoring of nearly the whole sky.
- Study transients with shorter time-scales than ever before, down to seconds.
- Potential to discover very rare
- and possibly yet unknown phenomena.
- Implications for astrophysics, cosmology and fundamental physics.



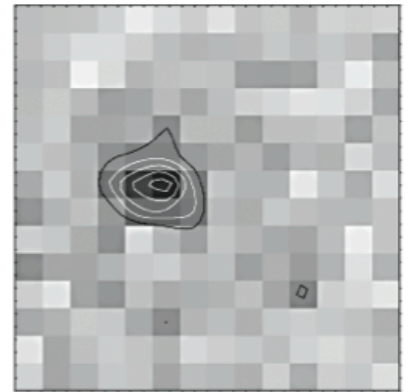
SN + Galaxy

—



Galaxy

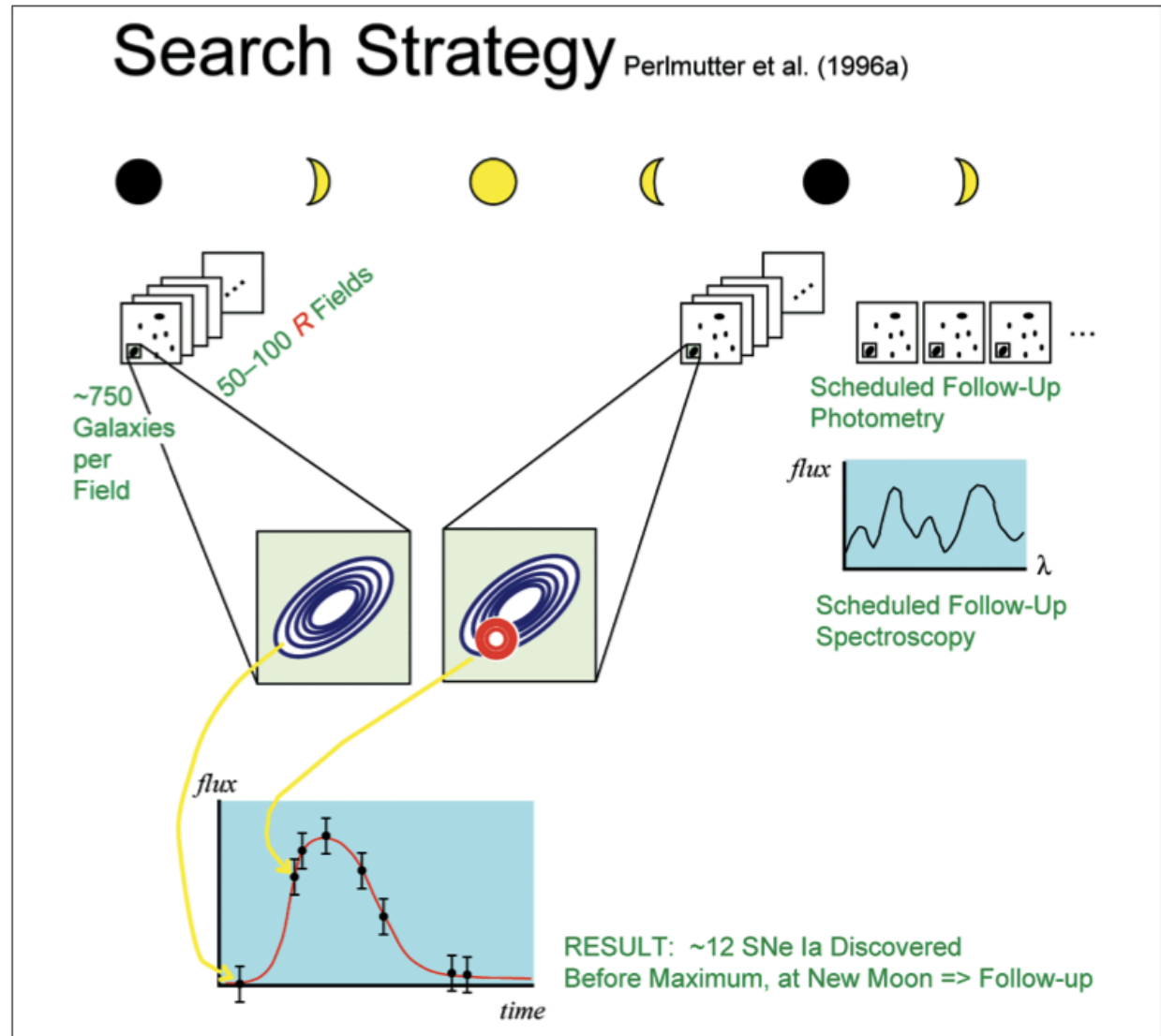
=



SN

SN Ia searches in the 90's and the discovery of dark energy

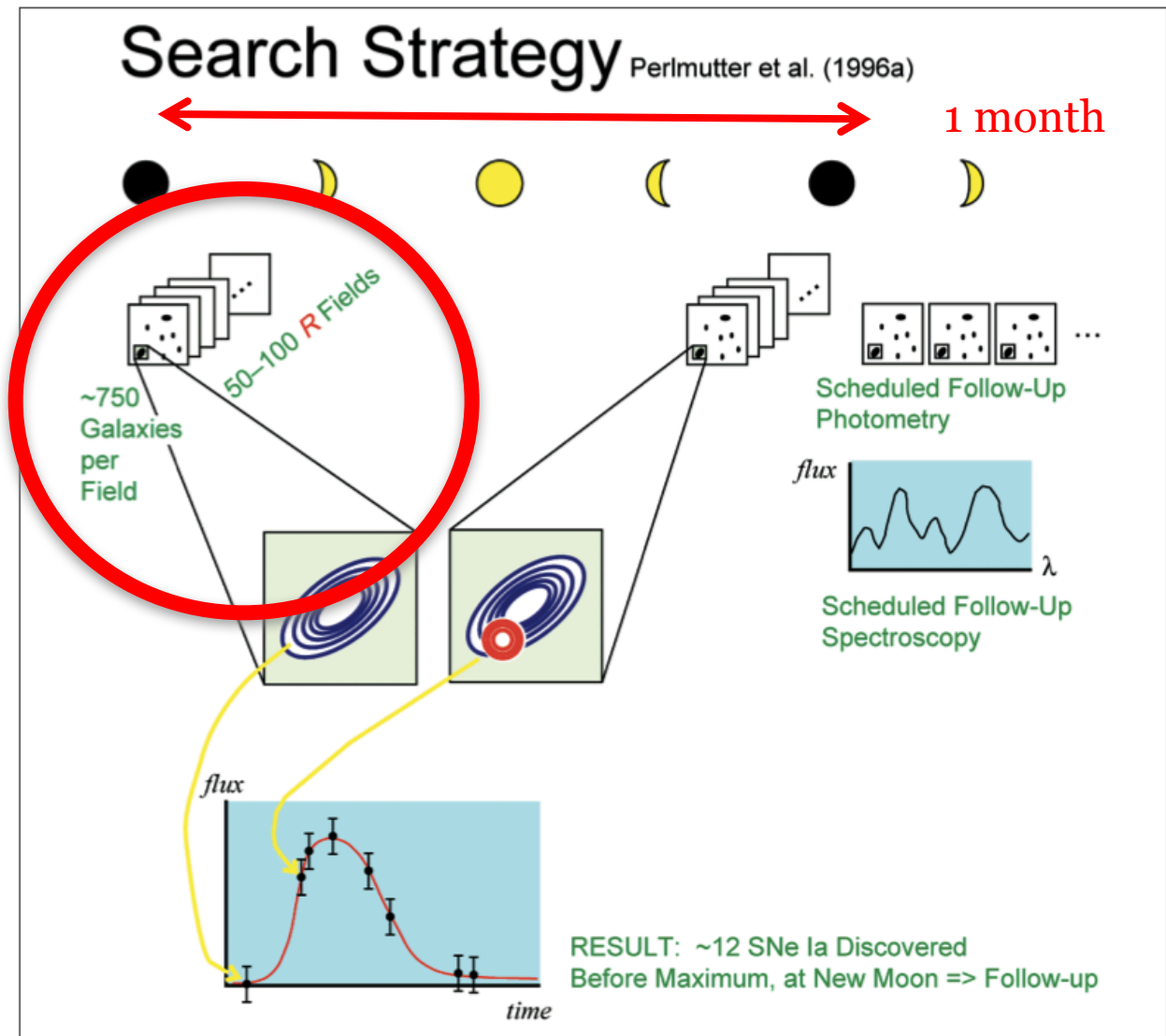
4



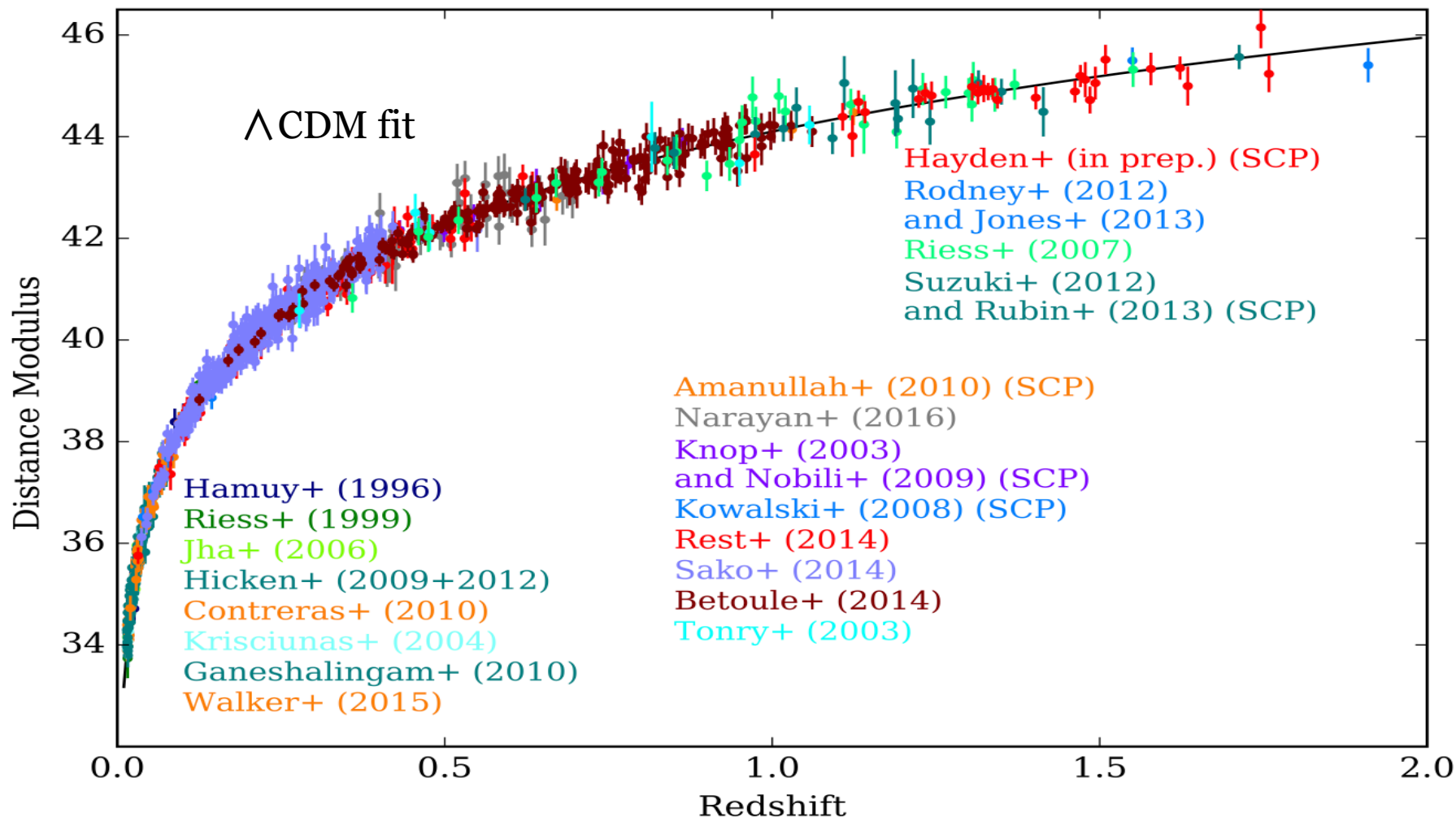
From Saul's Nobel Lecture (2011)

SN Ia searches in the 90's and the discovery of dark energy

5

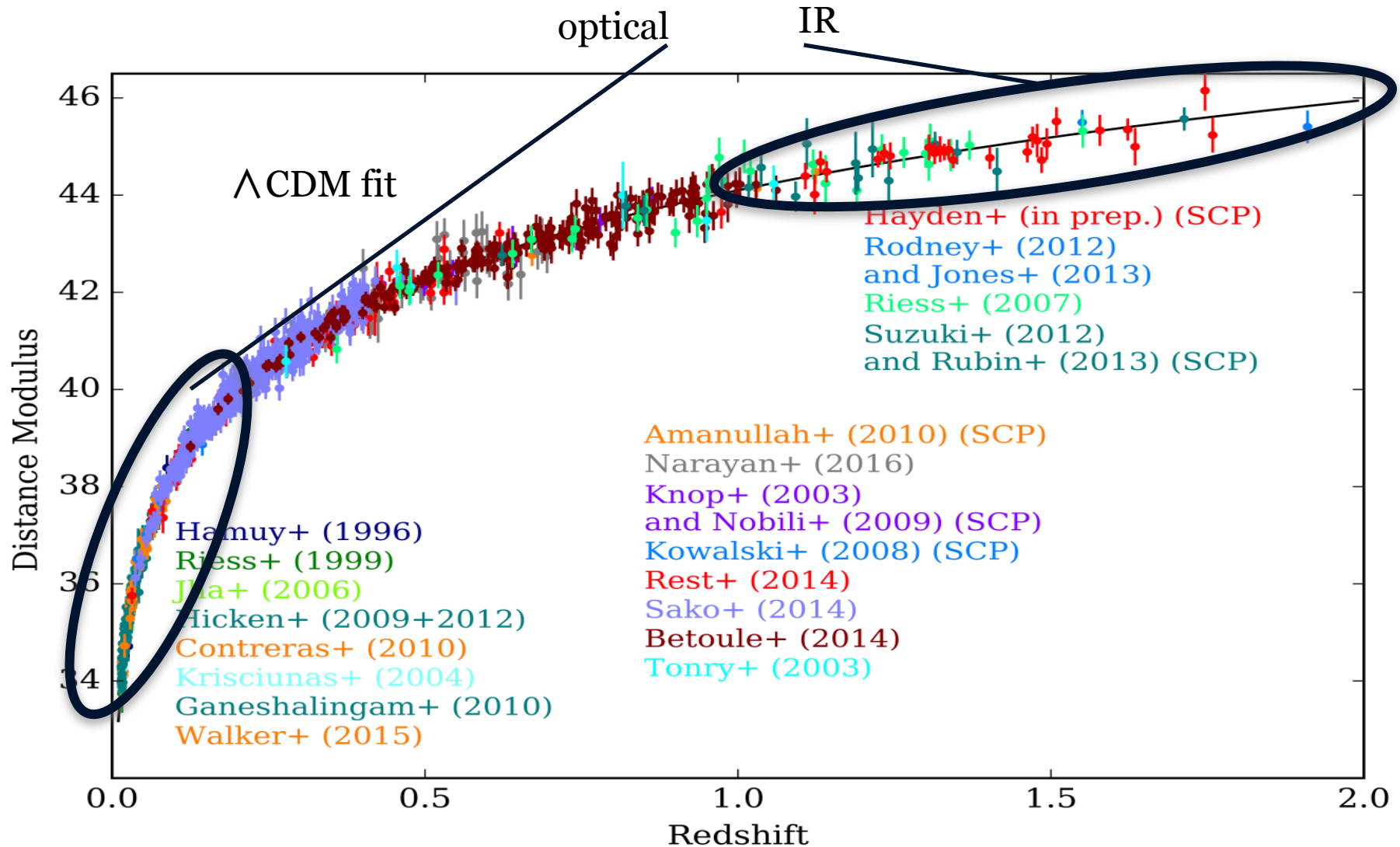


SN Ia Hubble diagram: v2017.9



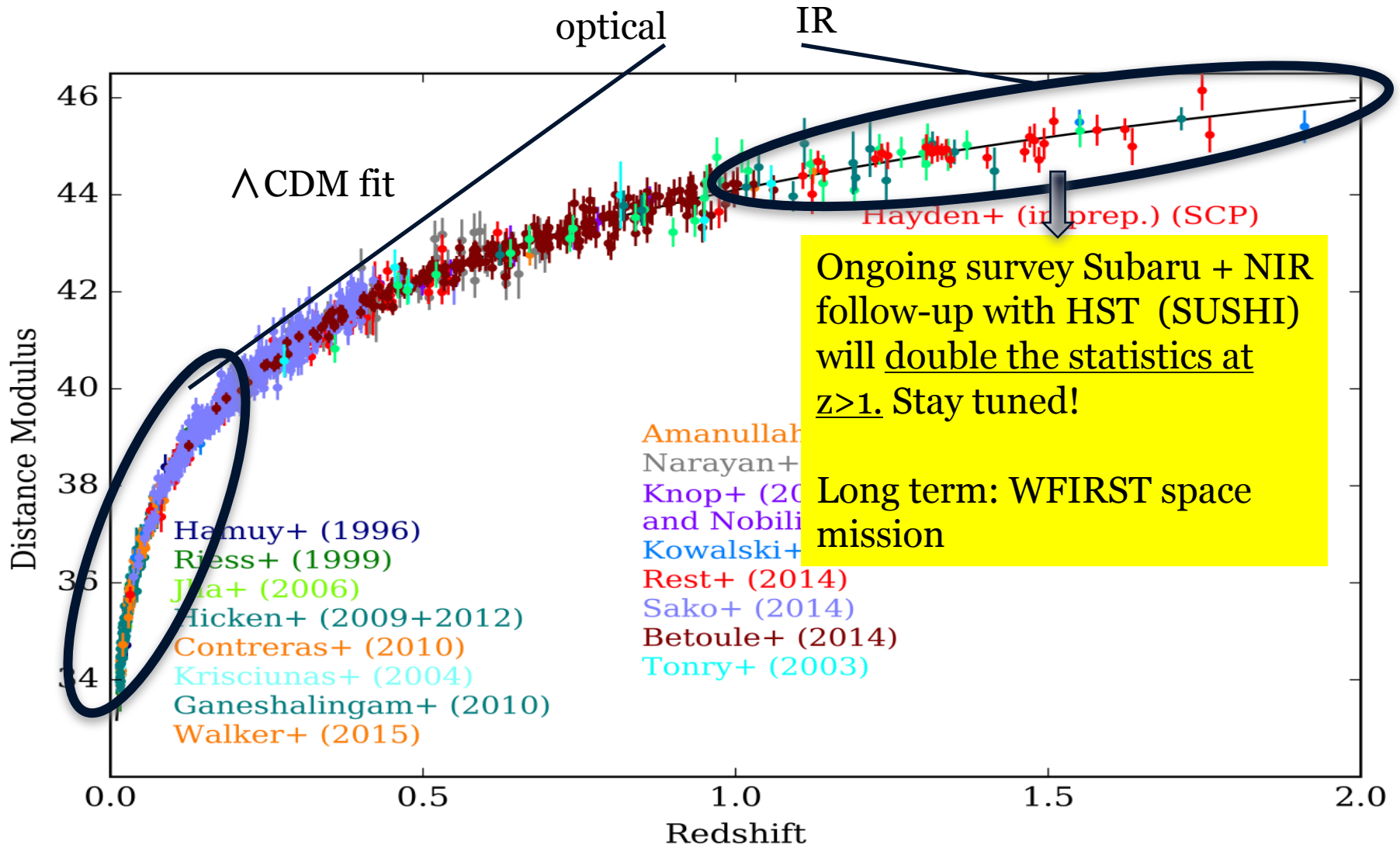
Further progress requires wide-field surveys

7



Further progress requires wide-field surveys

8



Important technological improvements

Larger CCD arrays: Field-of-view of cameras has changed from just a few sq.*arcminutes* to many sq.*degrees*

Large statistics and improved chances to catch very rare transient phenomena.

(Example will follow shortly)

Other implications:

High “cadence”, i.e., frequency by which the same piece of sky is revisited:

1) ability to discover transients/supernovae **much earlier** in the lightcurve and trigger follow-up observations with specialized instruments (e.g., from space)

2) Possibility to **find new transient phenomena**, with much shorter time scales

iPalomar Transient Factory (2009-2013-2017)

P48
survey telescope

P60
classification
telescope

P200
Spectroscopy

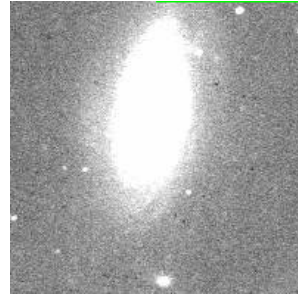
Caltech (PI: Shri Kulkarni)
University of Wisconsin
Los Alamos National Lab
Oskar Klein Centre, Stockholm
Weizmann Institute, Israel
Kavli@Tokyo
University of Taiwan
University of Maryland

100 Megapixel CCD
2.3 x 3.4 deg FOV

7.2 deg²
operational



Pipeline flow-chart



Raw data

(Palomar -> Caltech -> LBL)

processing

3 min

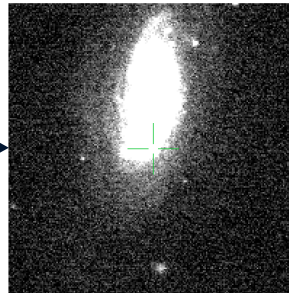
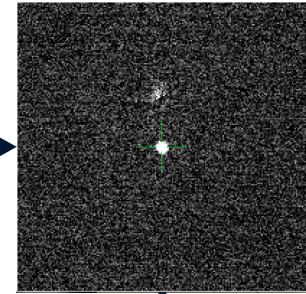


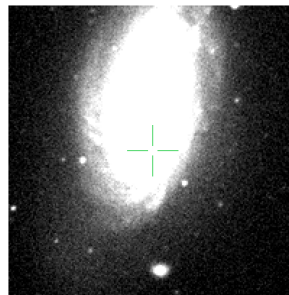
Image
Subtraction

1.5 min/
image



Real-Bogus
(reject false
alarms)
0.5 min/
image

Load
candidates
Into
database



Reference
Image

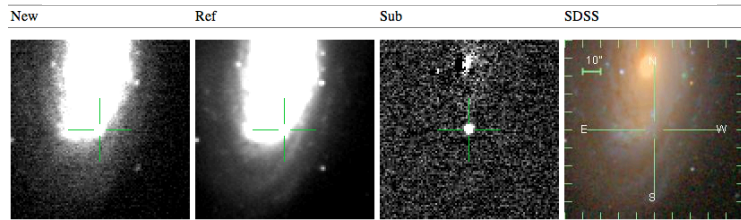
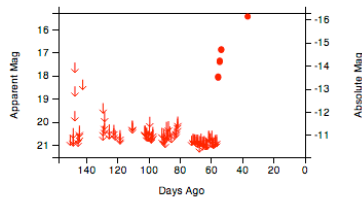
Star/asteroid
Identification
Local galaxy match

0.5 min/image



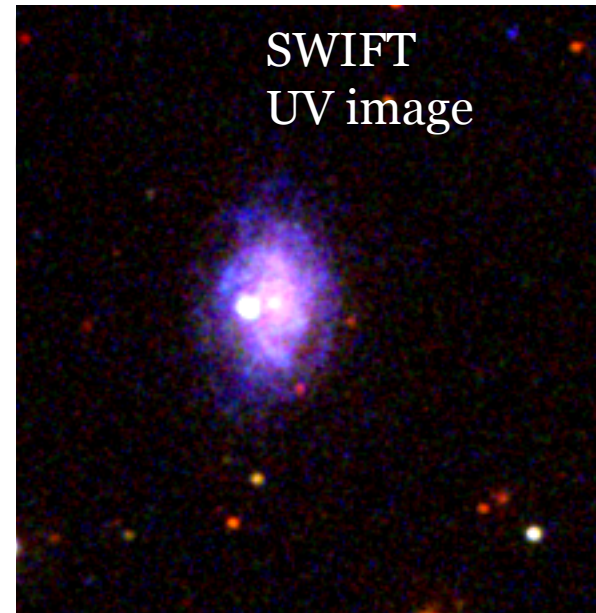
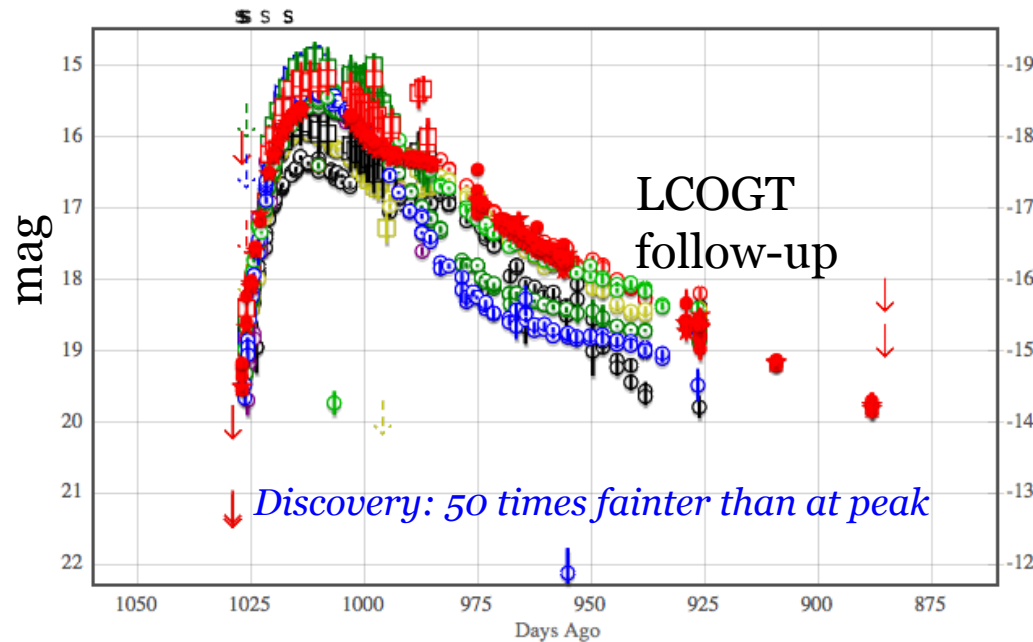
Database

ID: 66468689 [Examine](#), 226973 [Zoom-Sub](#)
RB2: 0.64
Mag: 17.33
iPTF [13bvn](#)
Nearby [PGC53578](#)
Abs Mag: -14.24
7 Matches in iPTF DB before tonight
0 Matches in PTF/best DB
Not a bad sub. 0.007, 0.191, 0.365



Scanning Page: human screening
mainly in Israel and Stockholm

COURTESY OF YI CAO & MANSI
KASLIWAL



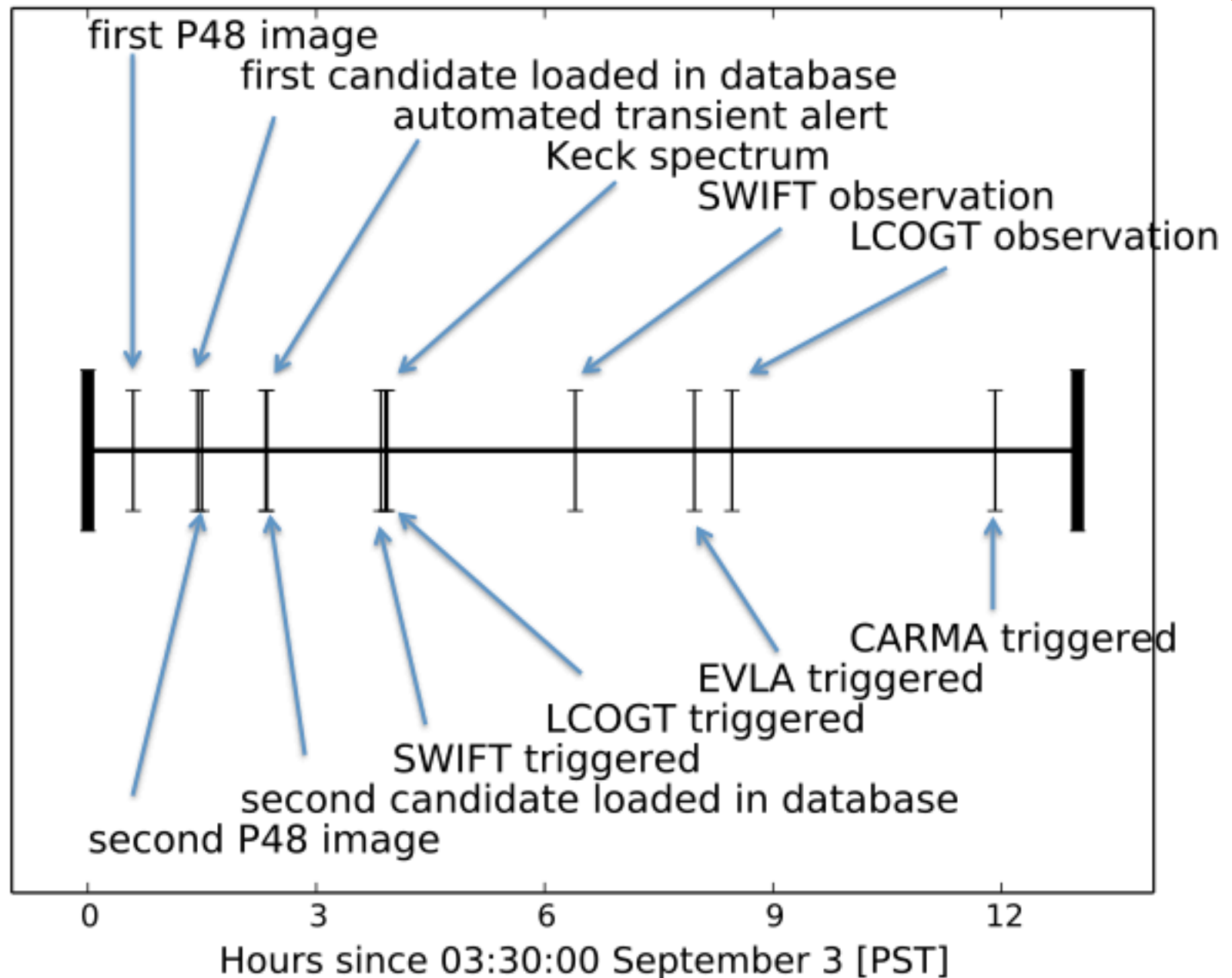
F481-band (Nugent LBVC pipeline v2)

Early observations of SNe – astrophysical unknowns:

- Search for evidence of shock heating of outer layers of exploding star, interaction with companion star, circumstellar medium, surface radioactivity, etc
- Multi-wavelength follow-up observations to study extinction along line of sight.

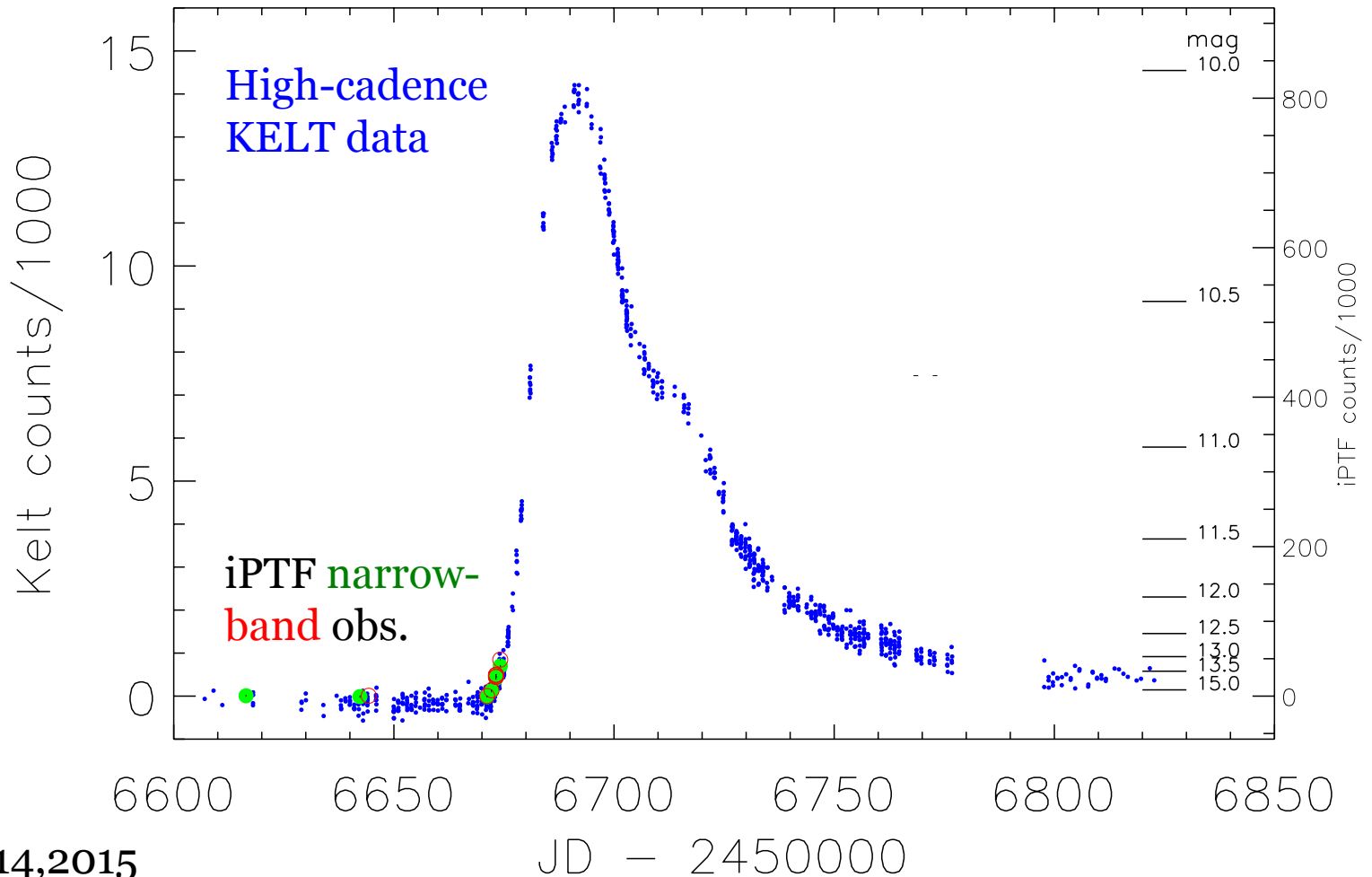
Time-line of follow-up of iPTF13dge

15



SN2014J in M82: see explosion *as it happens!*

16

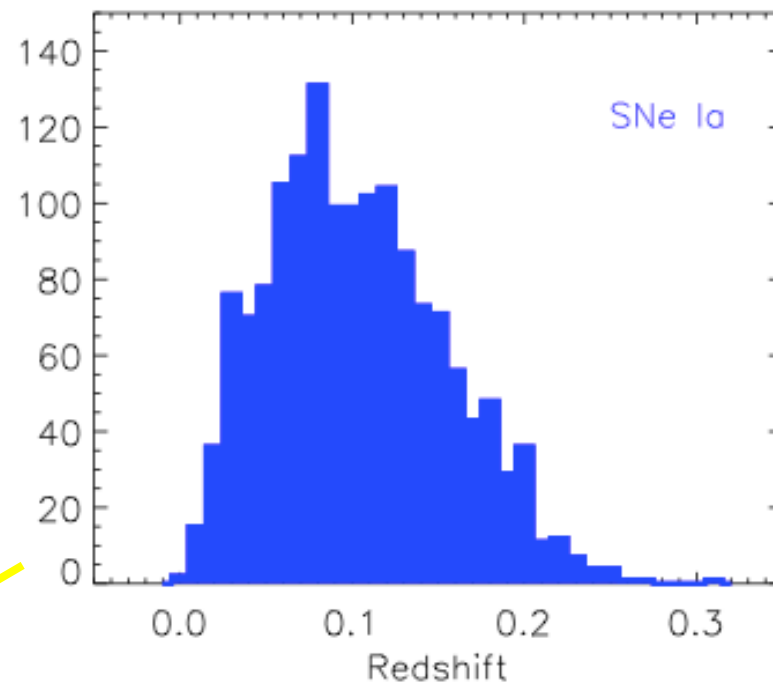
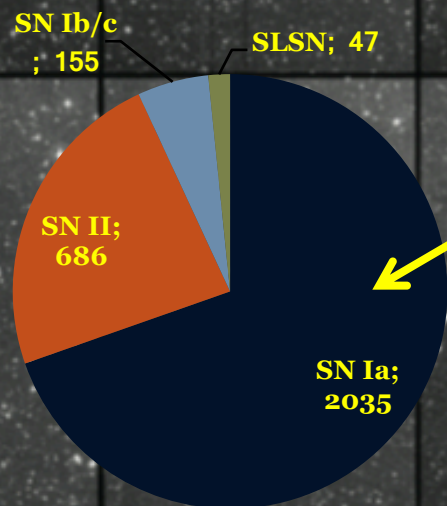


AG+2014,2015
 Siverd, AG+2015

Lots of follow-up observations/analysis!

100 Megapixel CCD
2.3 x 3.4 deg FOV

7.2 deg²
operational



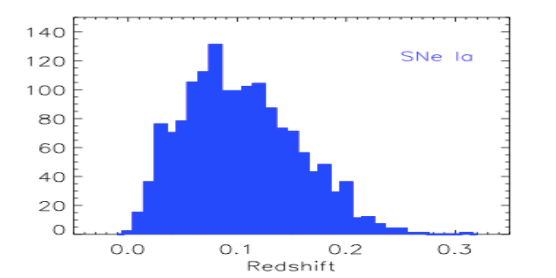
Typical operation: 60 s exposures + 36 s readout:
Pixel size: 1", S/N=5 for 21 mag in R and g



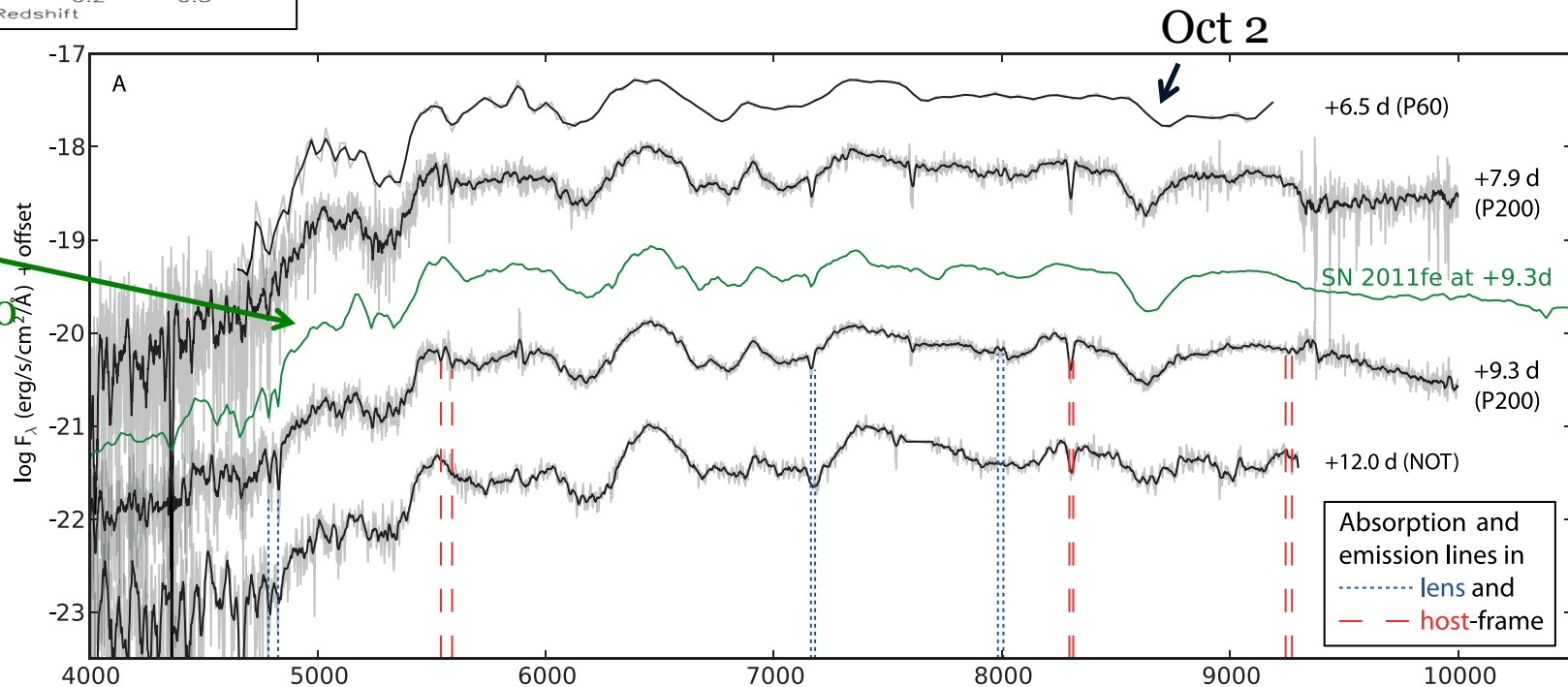
On Oct 2, 2016, something quite unexpected happened...

Perfect match to $z=0.409$ SN Ia

19



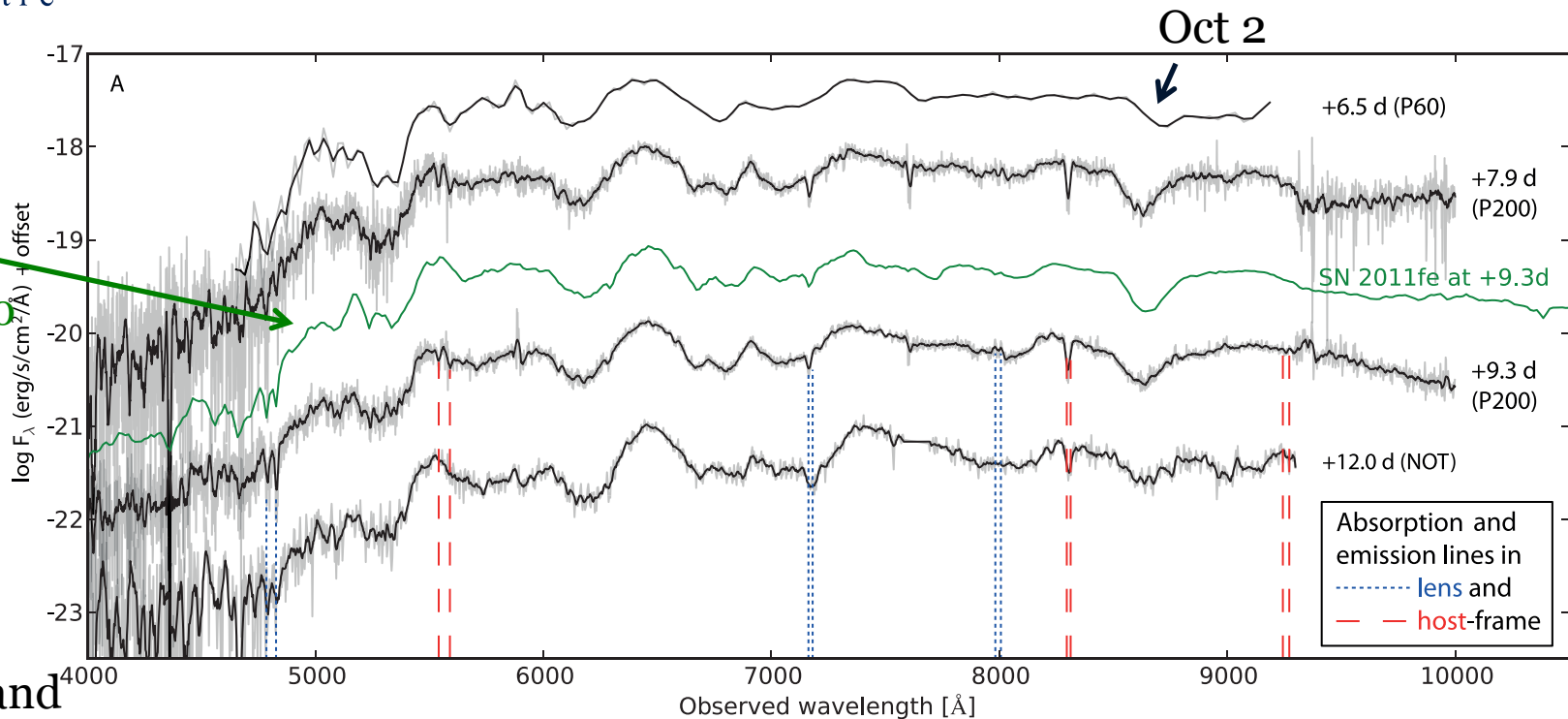
“Typical”
SNIa
redshifted to
 $z=0.409$



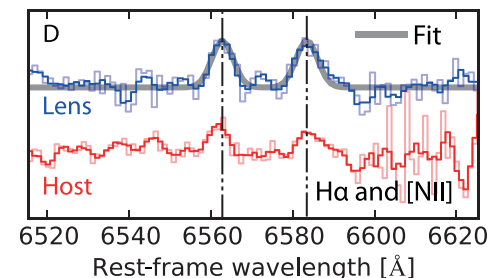
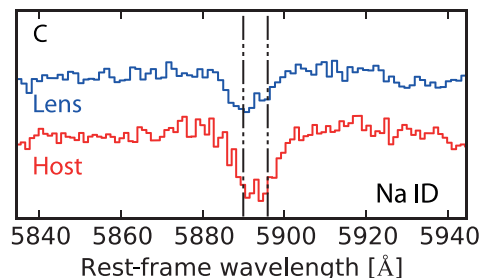
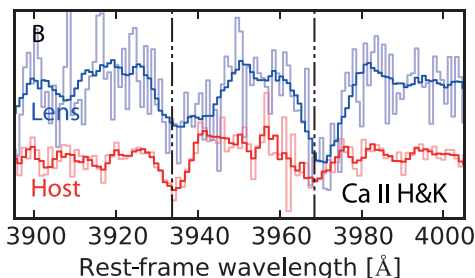
Perfect match to $z=0.409$ SN Ia + intervening galaxy at $z=0.216$

20

“Typical”
SNIa
redshifted to
 $z=0.409$



Absorption and
emission lines
at two redshifts:
the SN host
galaxy and
other galaxy in
line-of-sight



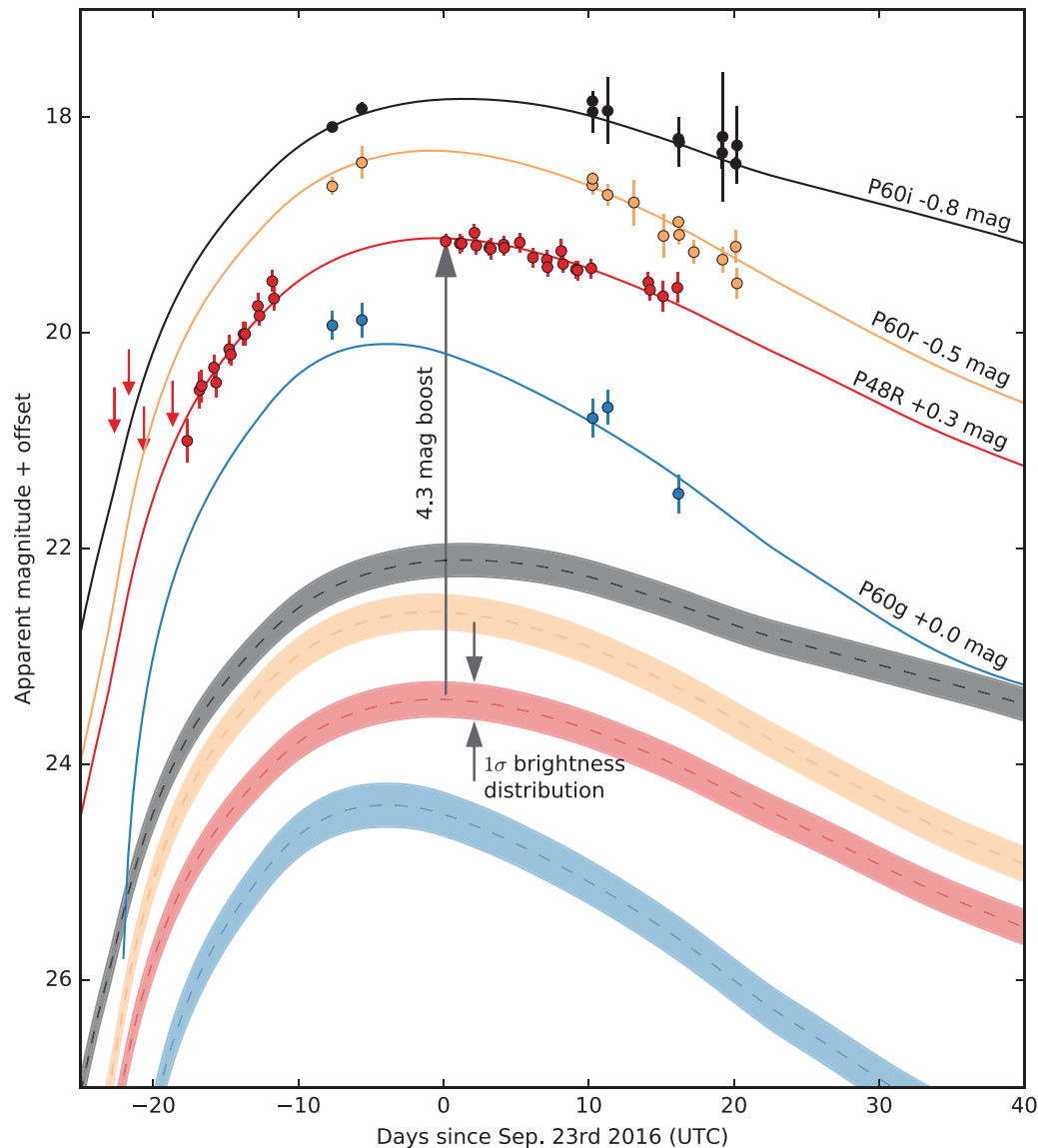
~52 times brighter than normal
SNIa at $z \sim 0.4$: a 30σ outlier!

21

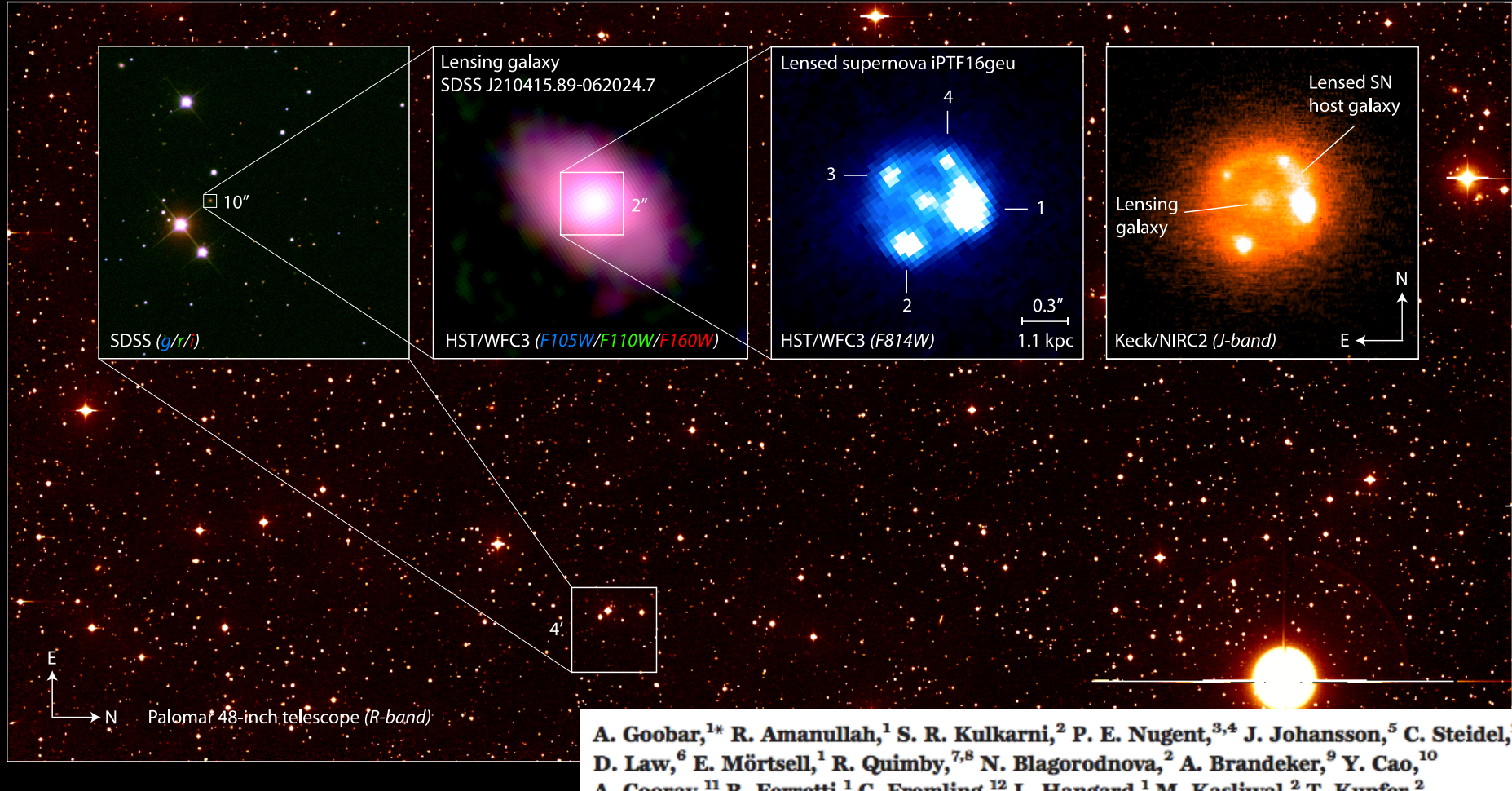
Only possible with
SNe Ia:
“standard candles”

Must be lensed by
the intervening
galaxy!

High-resolution
imaging with
HST requested +
Adaptive Optics
from VLT and
Keck.



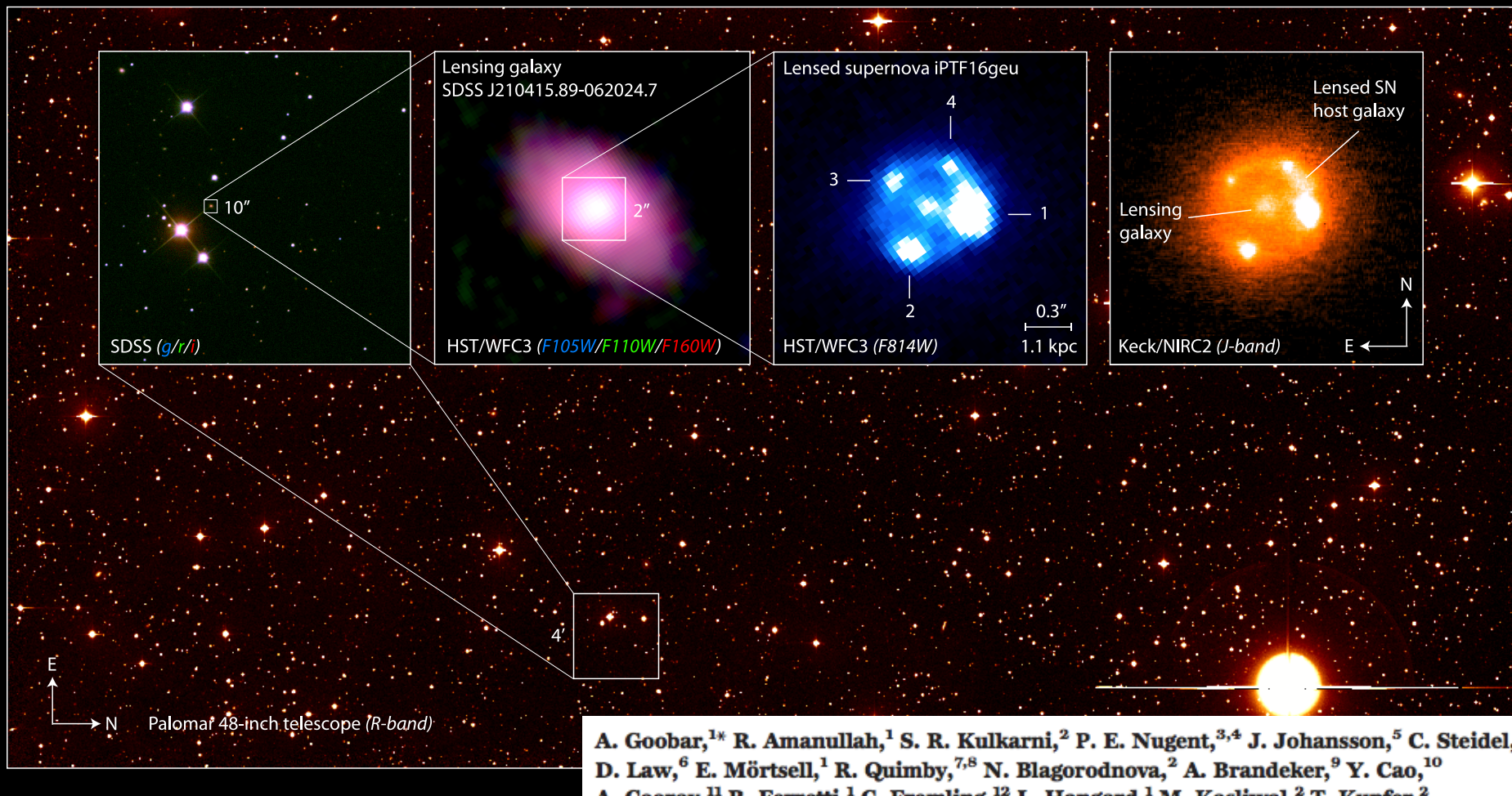
A quadrupole lens!



A. Goobar,^{1*} R. Amanullah,¹ S. R. Kulkarni,² P. E. Nugent,^{3,4} J. Johansson,⁵ C. Steidel,²
D. Law,⁶ E. Mörtzell,¹ R. Quimby,^{7,8} N. Blagorodnova,² A. Brandeker,⁹ Y. Cao,¹⁰
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R. Lunnan,^{2,9} F. Masci,¹³ A. A. Miller,^{14,15} H. Nayeri,¹¹ J. D. Neill,² E. O. Ofek,⁵
S. Papadogiannakis,¹ T. Petrushevska,¹ V. Ravi,² J. Sollerman,¹² M. Sullivan,¹⁶
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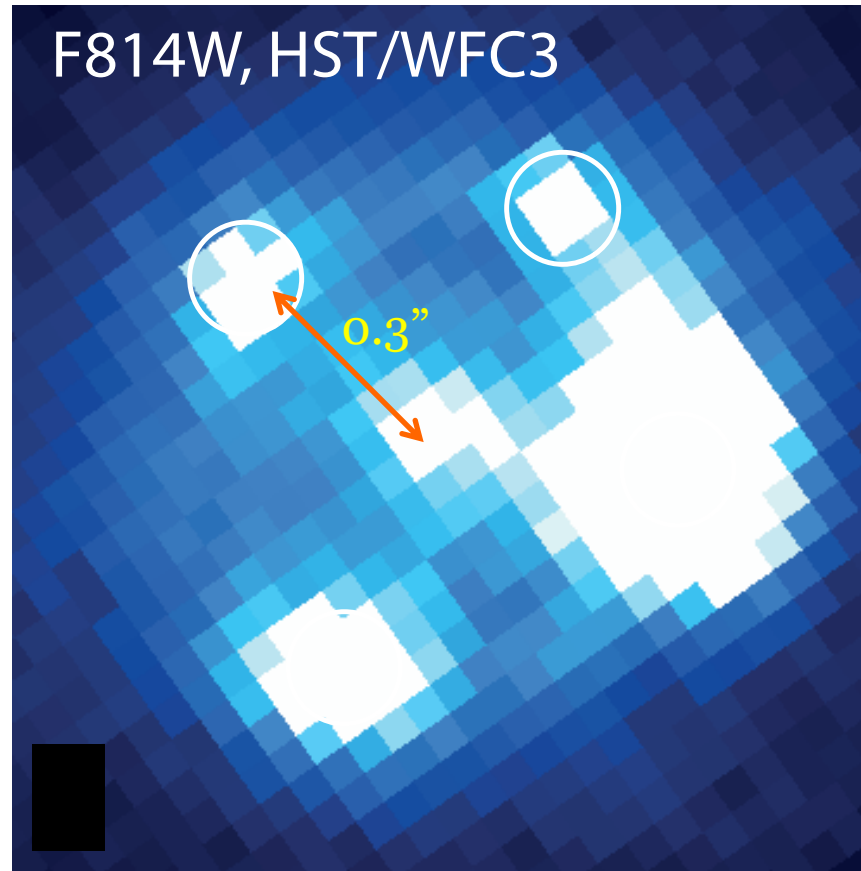
Time-delay measurement in progress... (H_0)

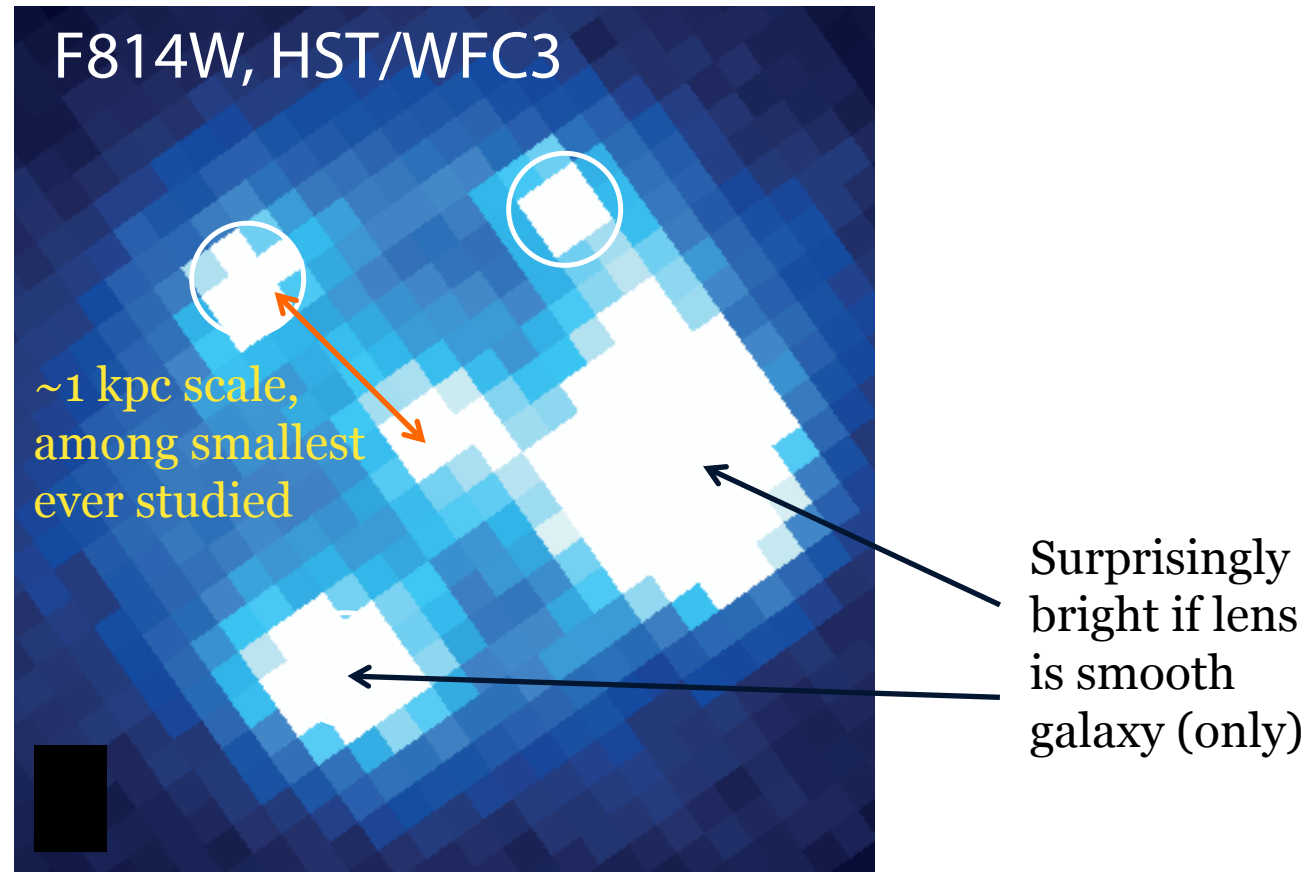


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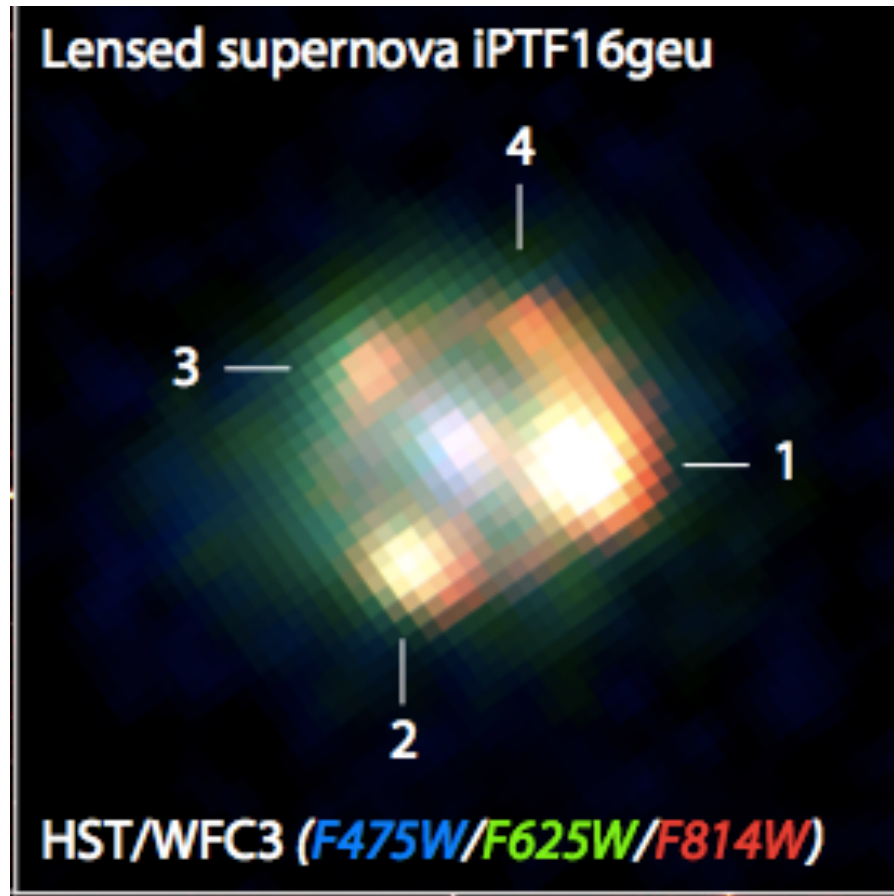
Discovery not limited by spatial resolution: it was the SN lightcurve that gave it away, 2" seeing!

24



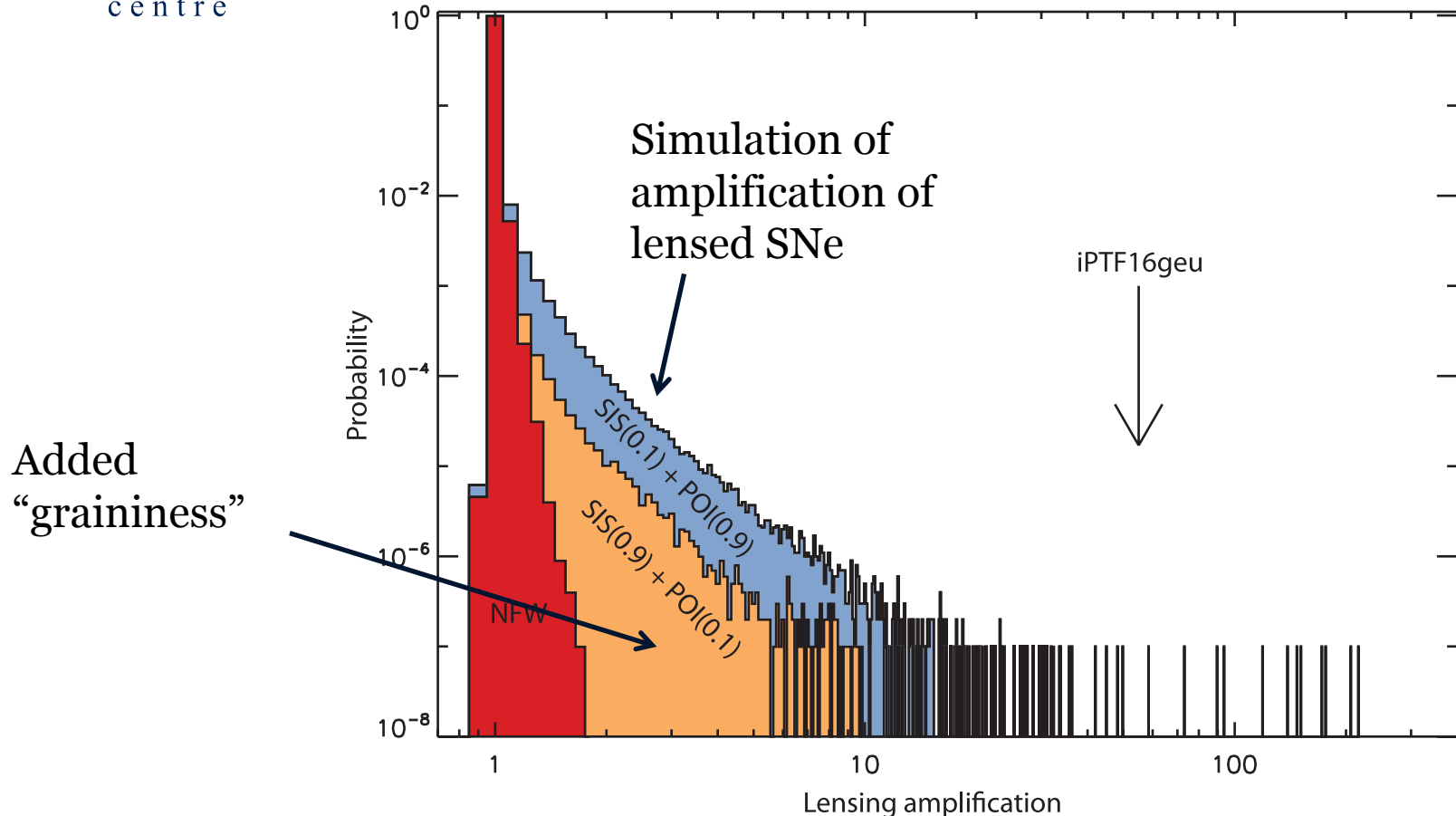


Large magnification and symmetry while very different intensities suggestive of added “compact” lenses in l-o-s

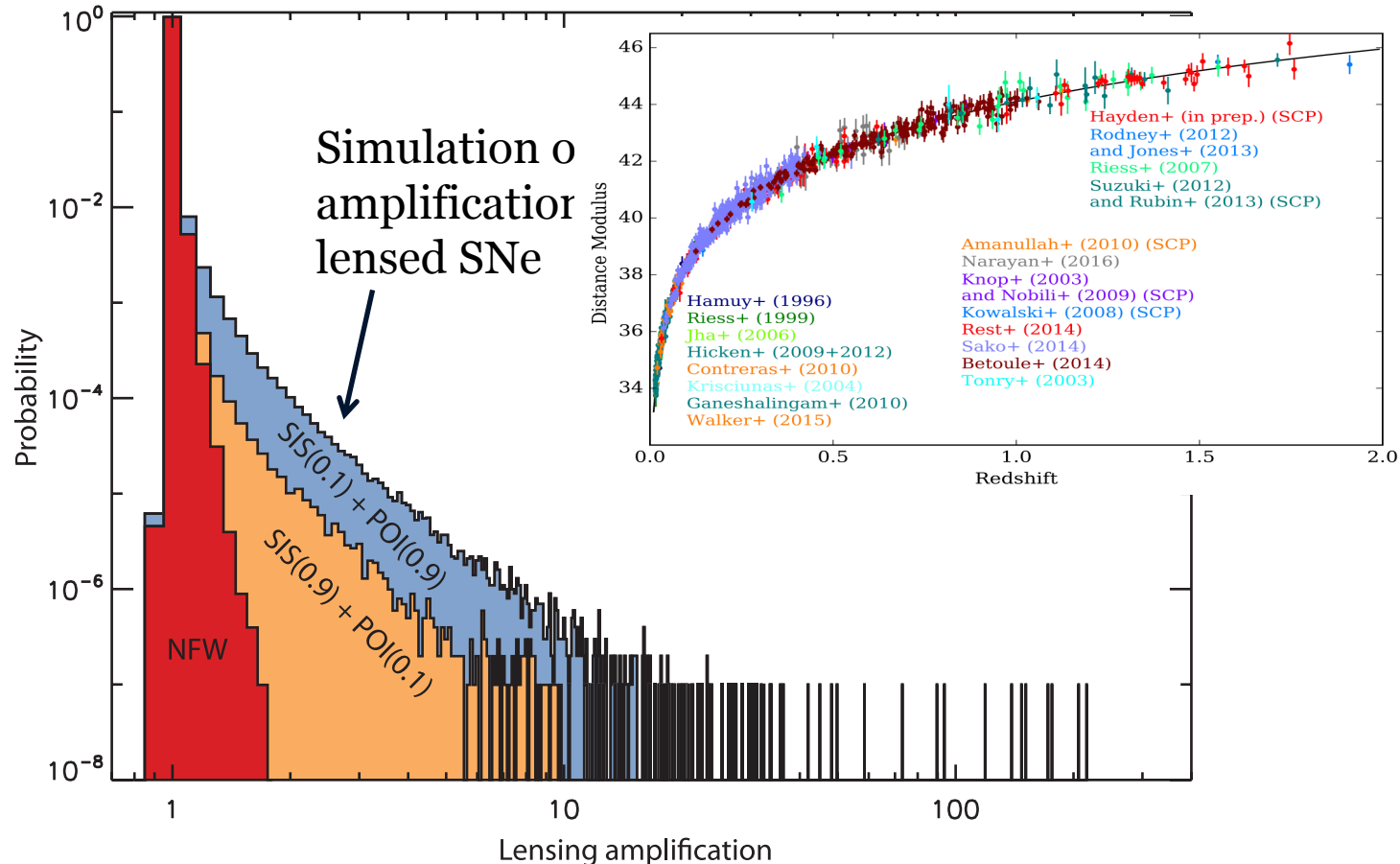


Evidence for reddening/extinction. But images 1 & 2 are $\sim 4\times$ brighter than expected compared to 3 & 4.

Large magnification and symmetry while very different intensities suggestive of added “compact” lenses in l-o-s



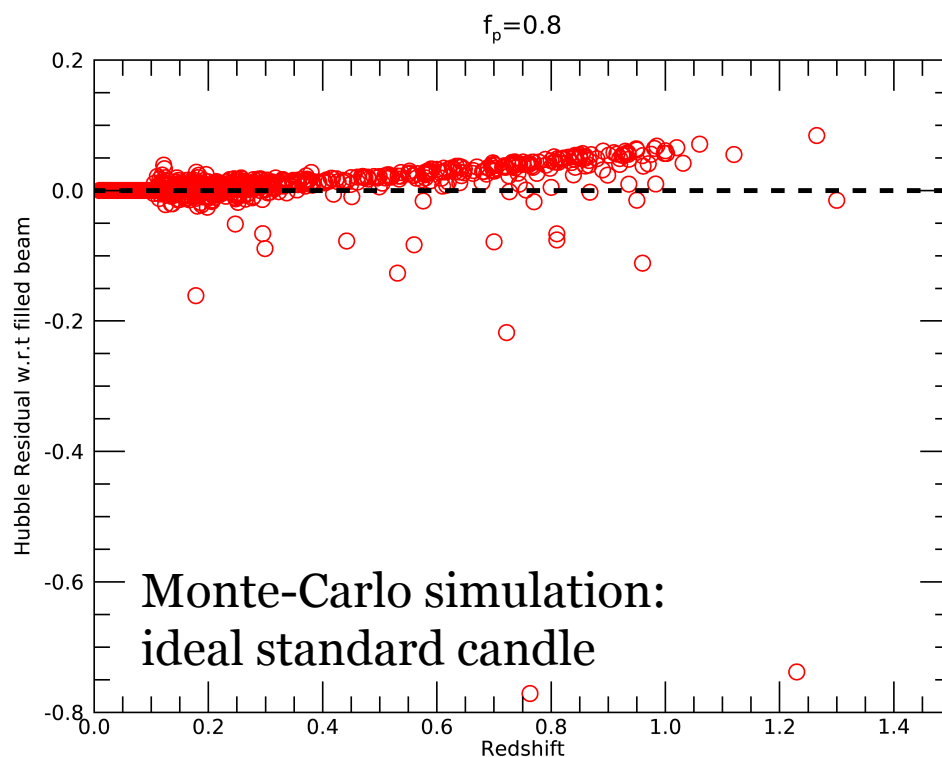
Measured amplification requires extraordinary alignment of source and lens – **also suggests lensing by substructures** in addition to a smooth matter component from the galaxy (e.g., network of stars, black holes, ...).
Just “lucky” or something profound?



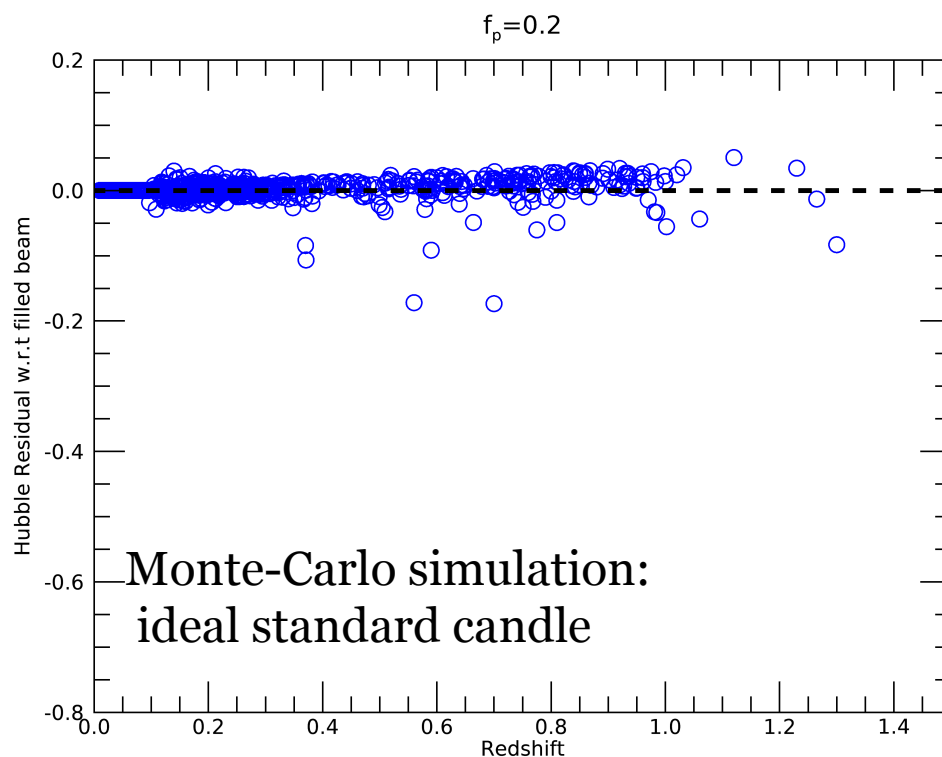
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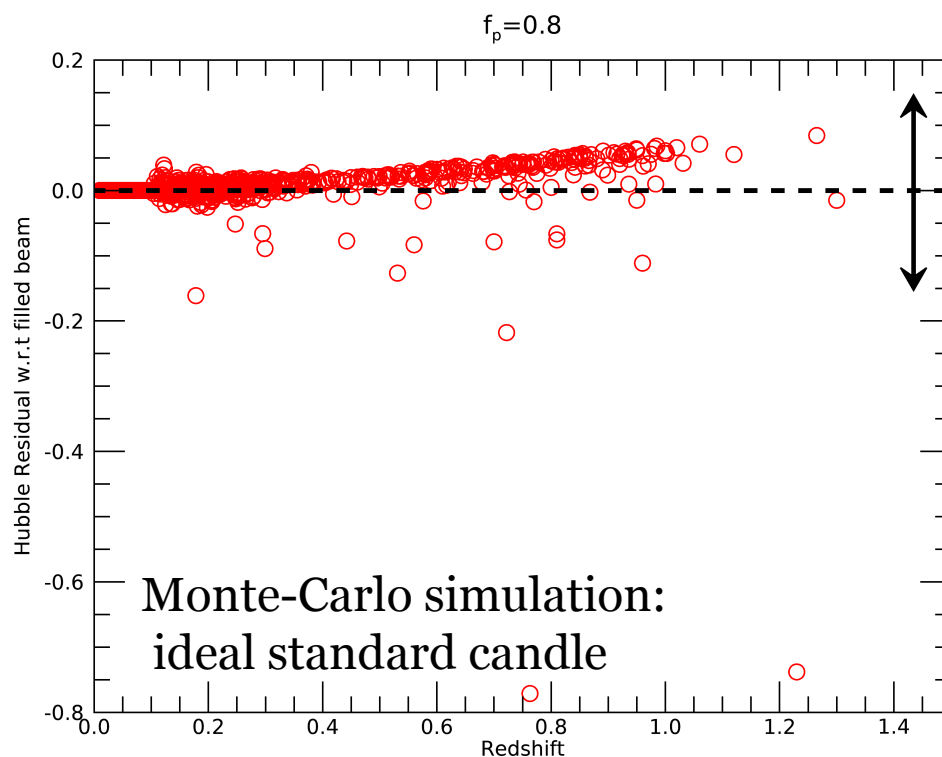
What can be said about clumping and inhomogeneities of DM from JLA sample?



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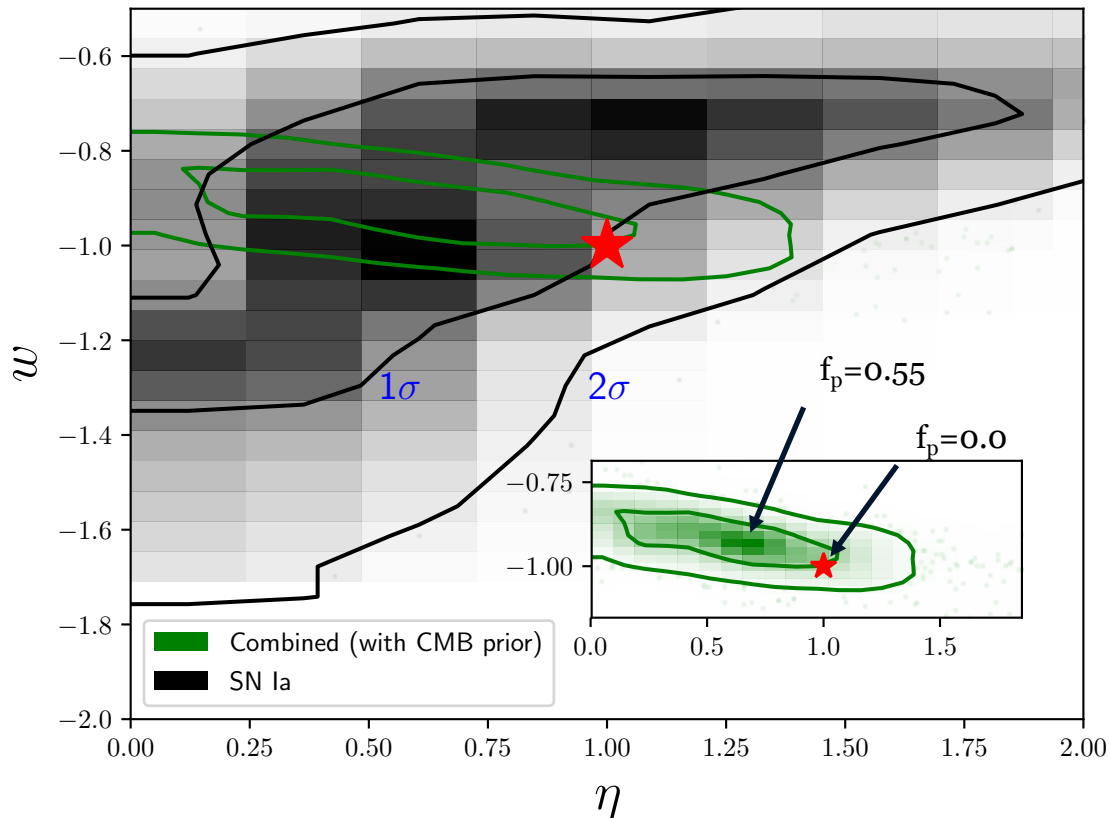


Typical measurement
error for 1 SN.

A quantitative study of JLA sample

- The distance-redshift **Dyer-Roeder** equation (Dyer & Roeder 1973) offers a way to quantify the level of inhomogeneous space-time probed by SN line of sight.
- The DR approximation assumes that the expansion rate of the Universe is governed by the total matter density whereas the focusing of light is only affected by a fraction η of the total matter density. $\eta=1$ implies homogeneous matter distribution ('filled beam').
- Fitting for η (+ all other cosmological parameters) shows **some degeneracy with Dark Energy equation of state, w .**

Bounds on inhomogeneous matter distribution



Planck + JLA SNIa

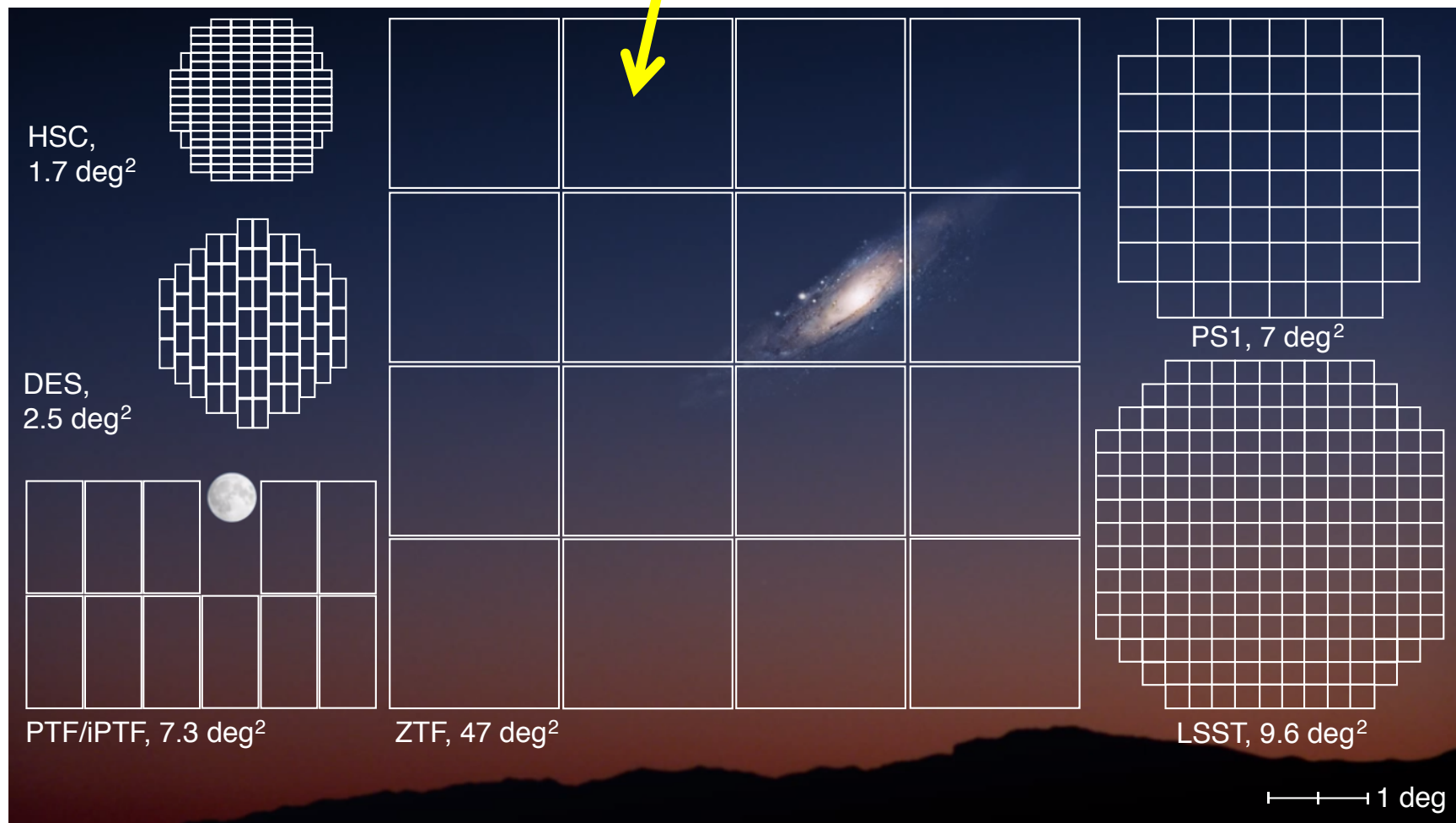
Current data **not yet sensitive enough**

Dhawan, AG, Mörtzell, *in prep*

New horizons in time-domain astrophysics.

- **Zwicky Transient Facility (ZTF)** is being installed @Palomar. Same telescope as iPTF, but much larger FoV and new camera. 3-year survey funded. Will collect all the low-hanging fruits at low redshifts. “Stepping stone” for LSST in TDA.
- **Large Synoptic Survey Telescope (LSST)** to see first light in 2020. 10-year “all sky” optical survey with suitable cadence for high- z SNe Ia.
- **WFIRST** (2025?) will operate at Near-IR wavelengths, sensitivity to detect first generation of cosmic explosions, besides doing “bread and butter” SNIa science.

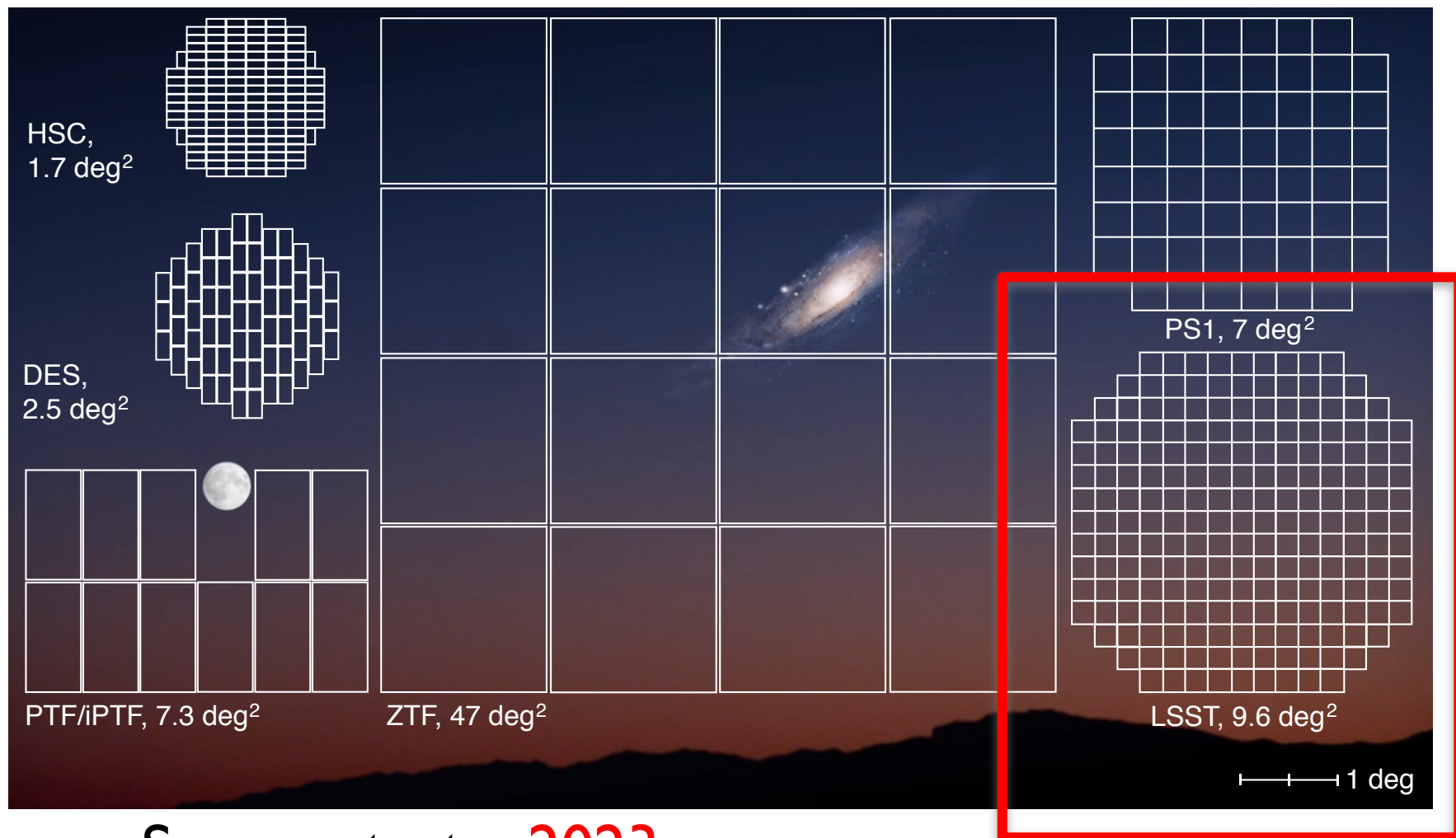
Zwicky Transient Facility: 35 12 x faster than iPTF!



First light: **one week from today!!**

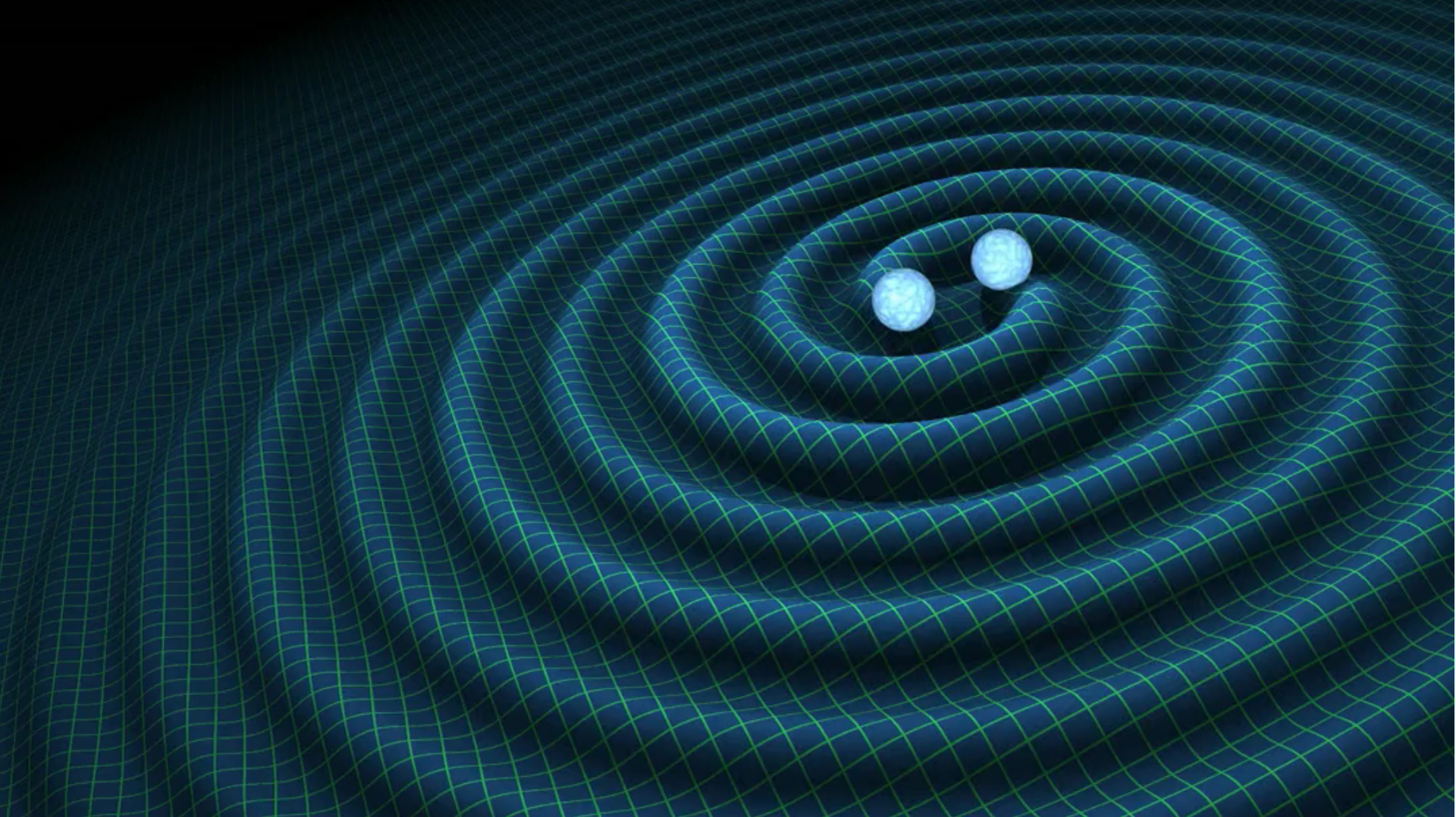
LSST: 8.4m mirror → substantially deeper

36



Survey start: ~2023

Next big leap? GW-EM counterparts?



GW-EM counterparts: “Macronovae”

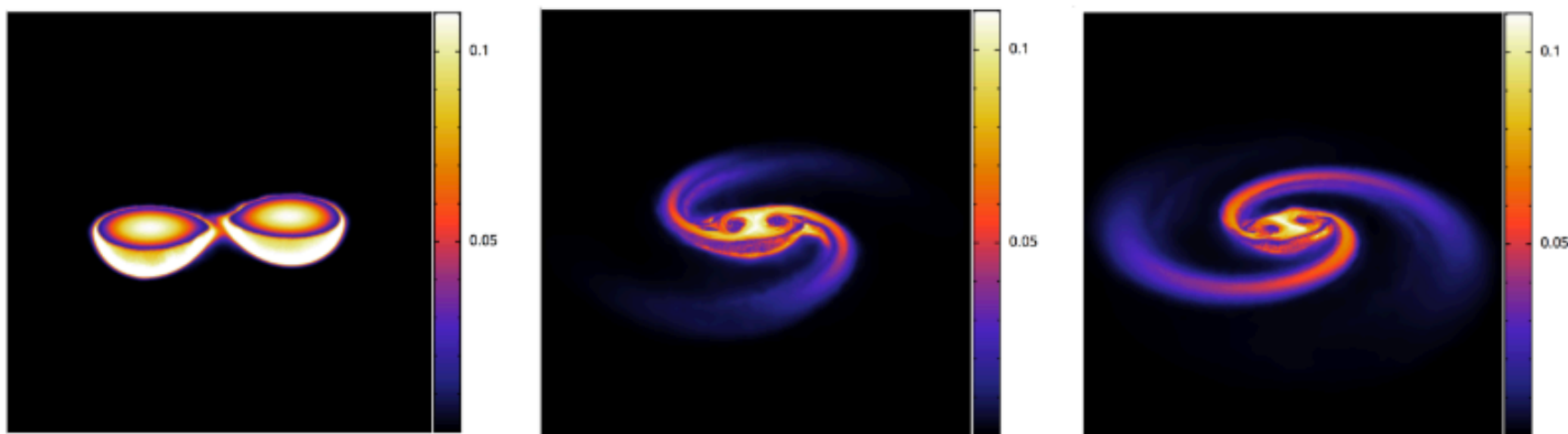
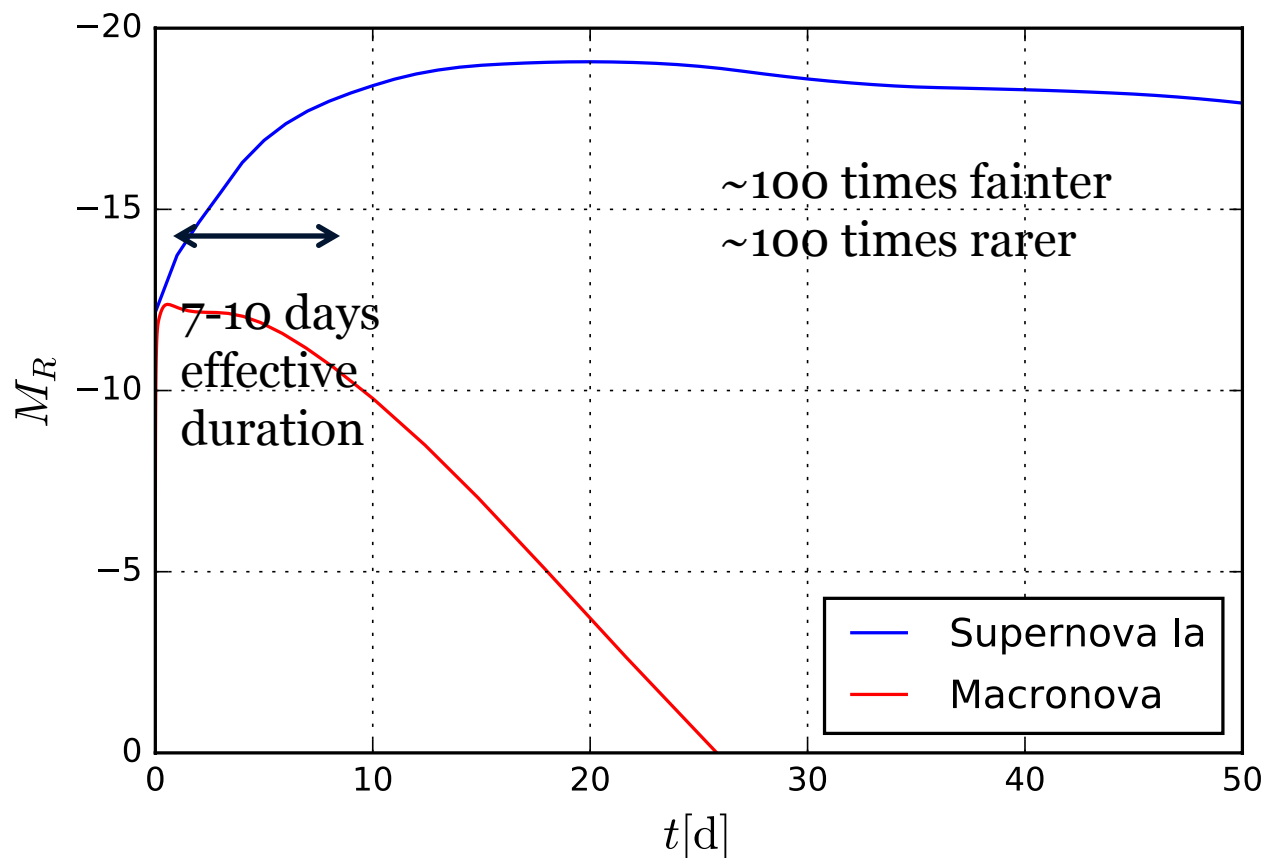


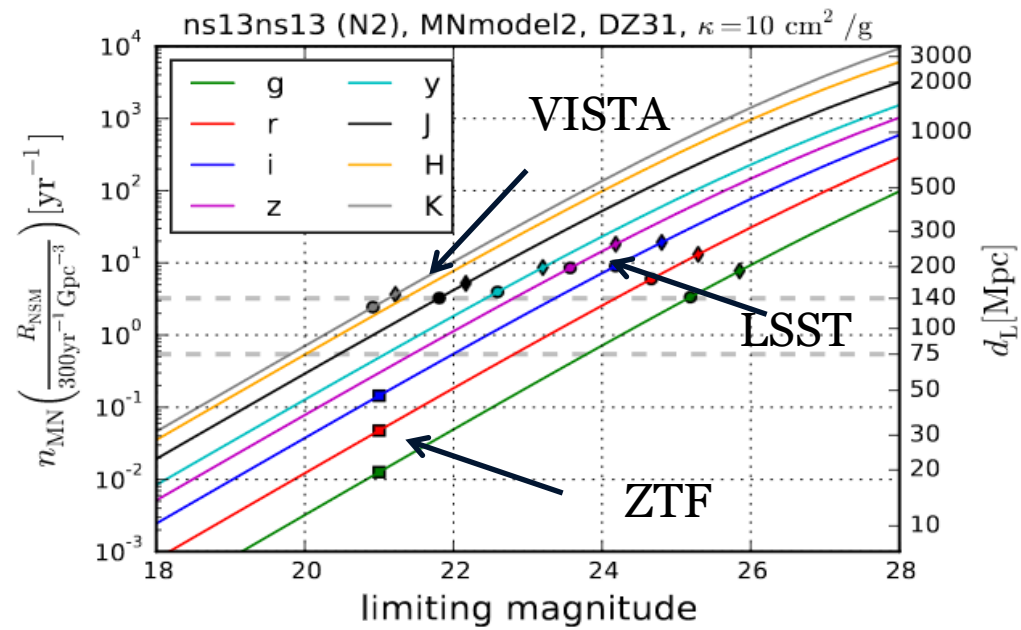
Figure 3. Electron fraction in a 1.3-1.3 M_{\odot} merger (model N2; only matter below orbital plane shown) at $t = 7.06, 11.6$ and 12.4 ms.

Merger of
neutron star- neutron star or
black hole- neutron star

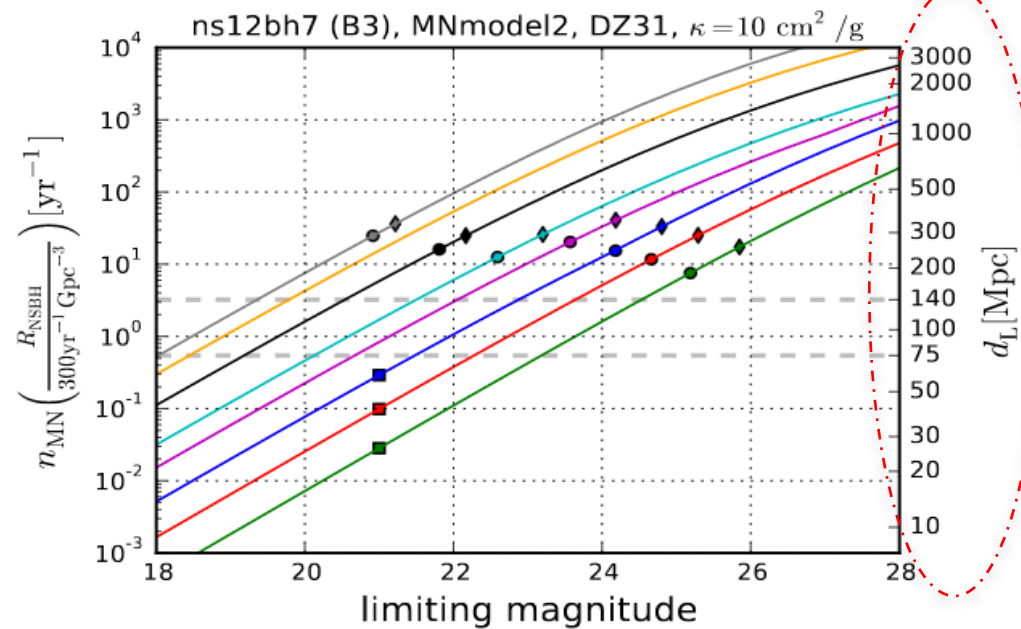
Rosswog et al 2017

Model prediction: Rare, Faint, Red and Fast!



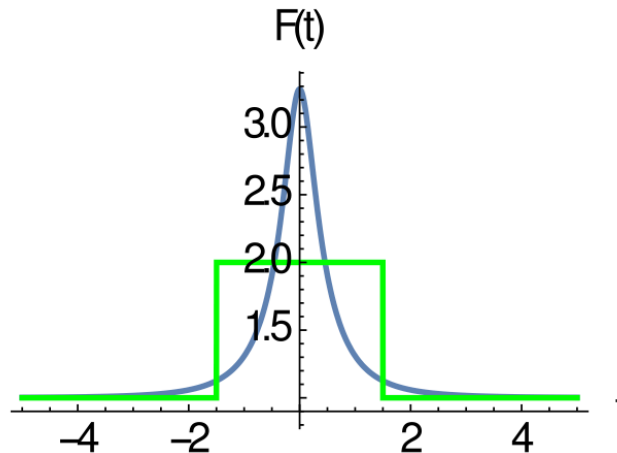
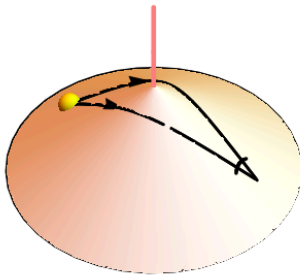


Rosswog et al 2017



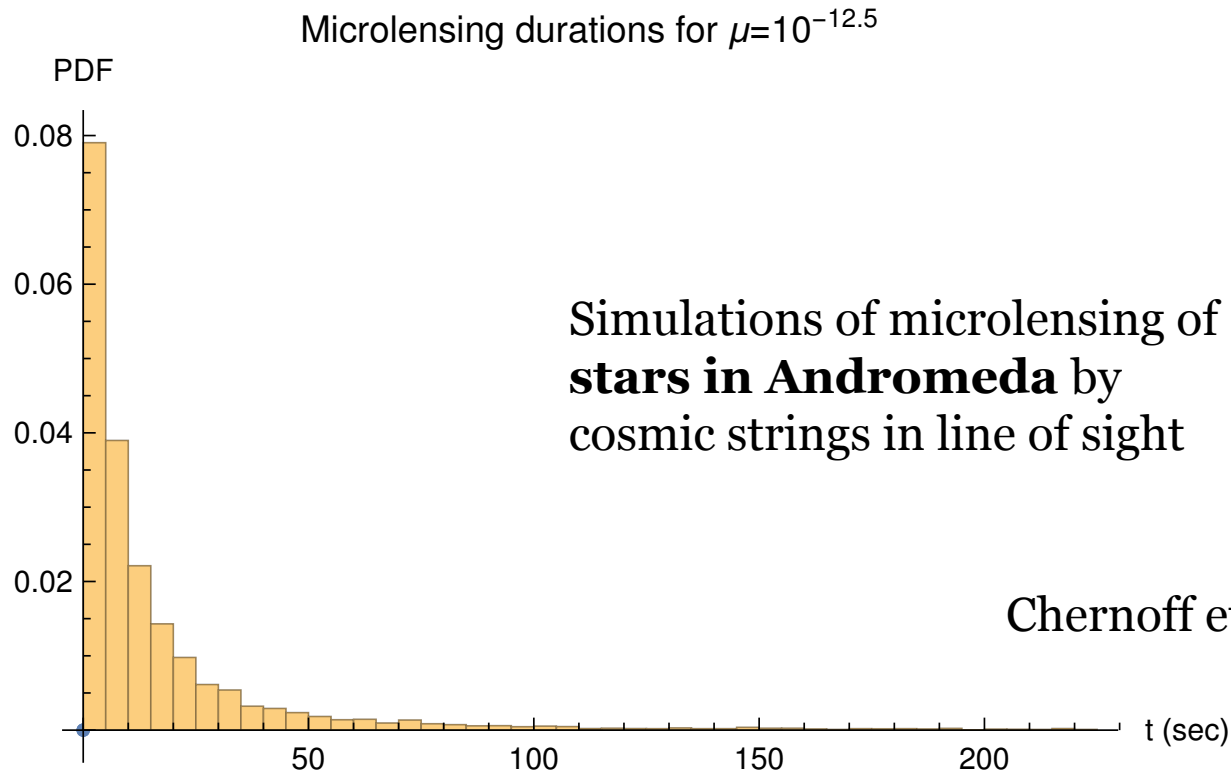
Exotic transients?

- **Microlensing** of stars by *early universe relics*. Very short time scales (**seconds or minutes!**) may be expected.
- **Cosmic Strings**, cosmological defects from Early Universe.



Chernoff & Tye

Short time scales: new frontier (I)

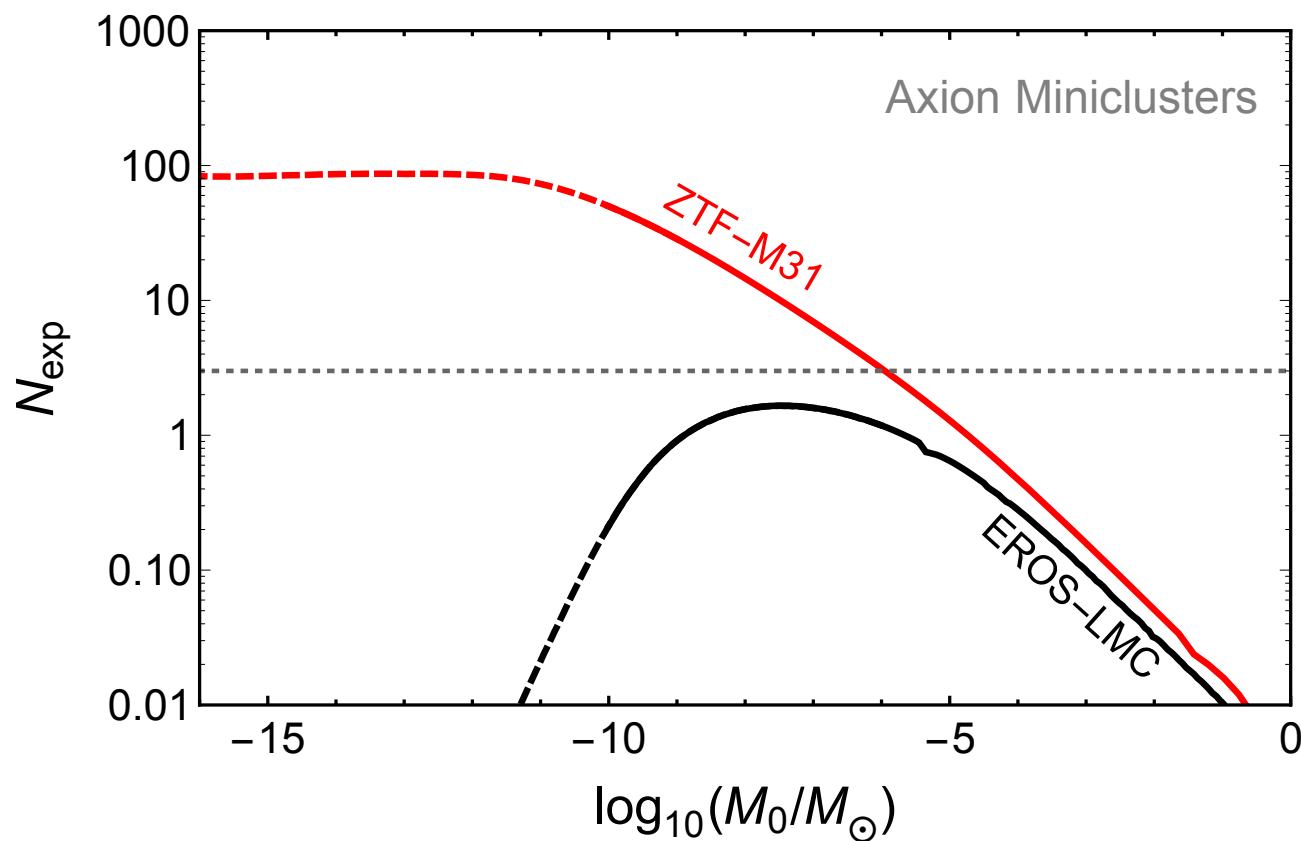


Chernoff et al, *in prep*

Time scale depends on string tension, μ

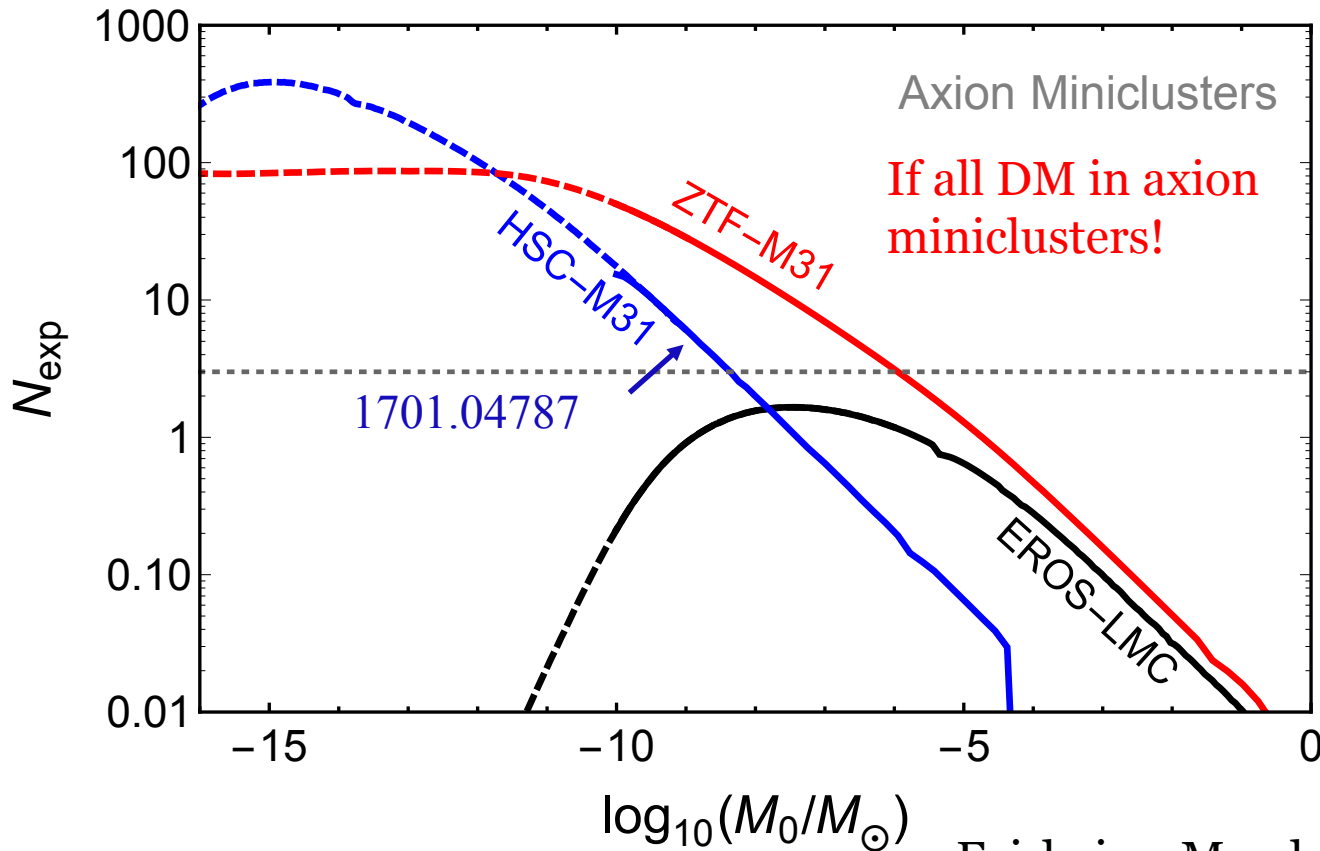
Short time scales: new frontier (II)

- Gravitational microlensing by **Axion “miniclusters”**
(+ **similar efficiency for light PBHs**)



Short time scales: new frontier (II)

- Gravitational microlensing by **Axion “miniclusters”**
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What's new?

- Monitoring of nearly the whole sky over much shorter time scales than ever before, down to seconds. *E.g., short time microlensing, onset of supernovae, anchor sample for SNIa cosmology + systematics.*
- Potential to discover very rare and possibly yet unknown phenomena. *E.g., gravitational lensing of standard candles in poorly explored length scales + H_0 ; GW-EM counterparts, microlensing by early universe relics, etc.*
- Bottle-neck? limited resources to follow-up the large number of discoveries. Have a telescope? Want to collaborate?

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- Bottle-neck? limited resources to follow-up the large number of discoveries. Have a telescope? Want to collaborate?
- **Please talk to me if you have other ideas for what we should be looking for with ZTF/LSST!**

Thank you!